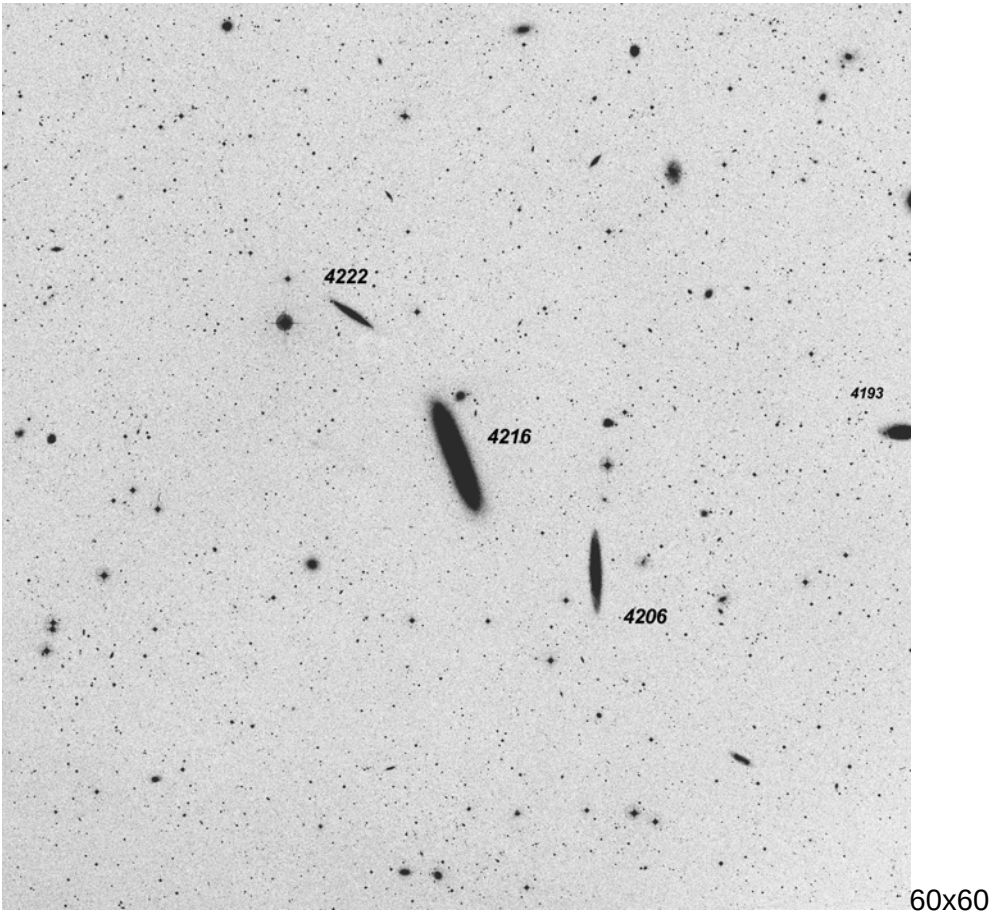


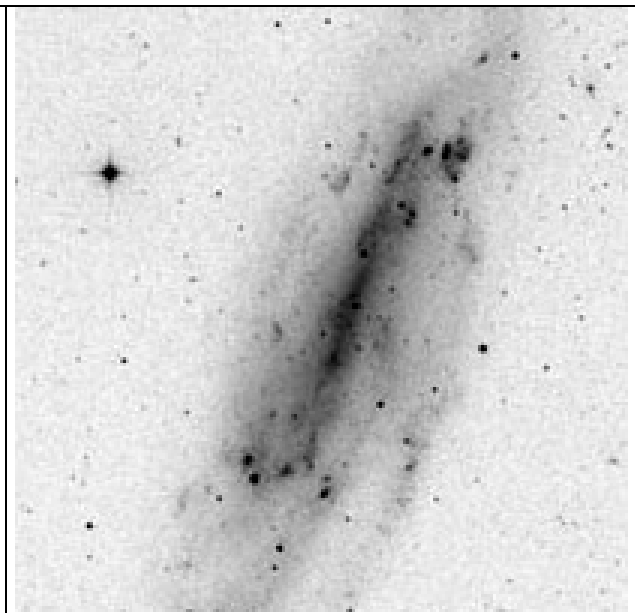
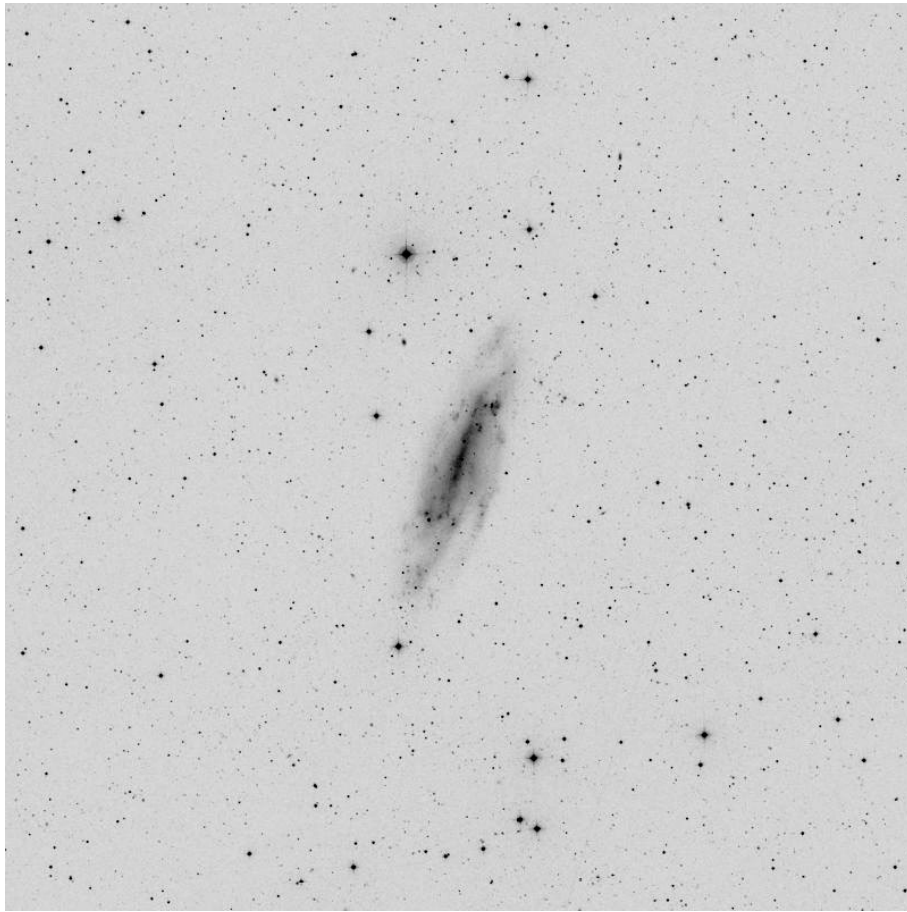
NGC 4216 and friends in Virgo



	other	RA	Dek	comments
NGC 4216		12 15 54	+13 08 58	

Observing notes:

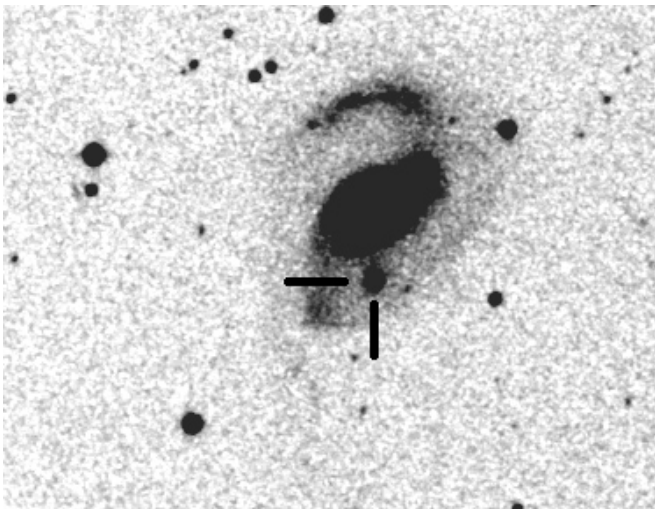
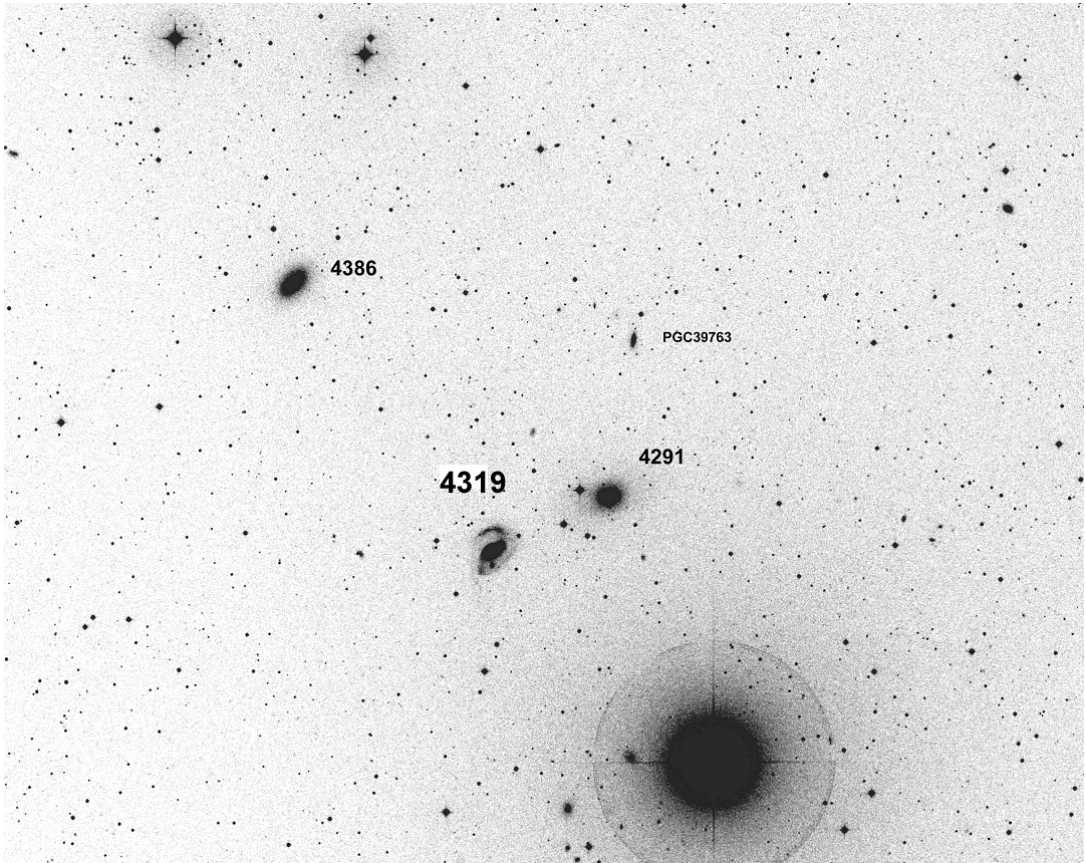
NGC 4236 in Draco



NGC 4236

12 16 41 +69 28 09

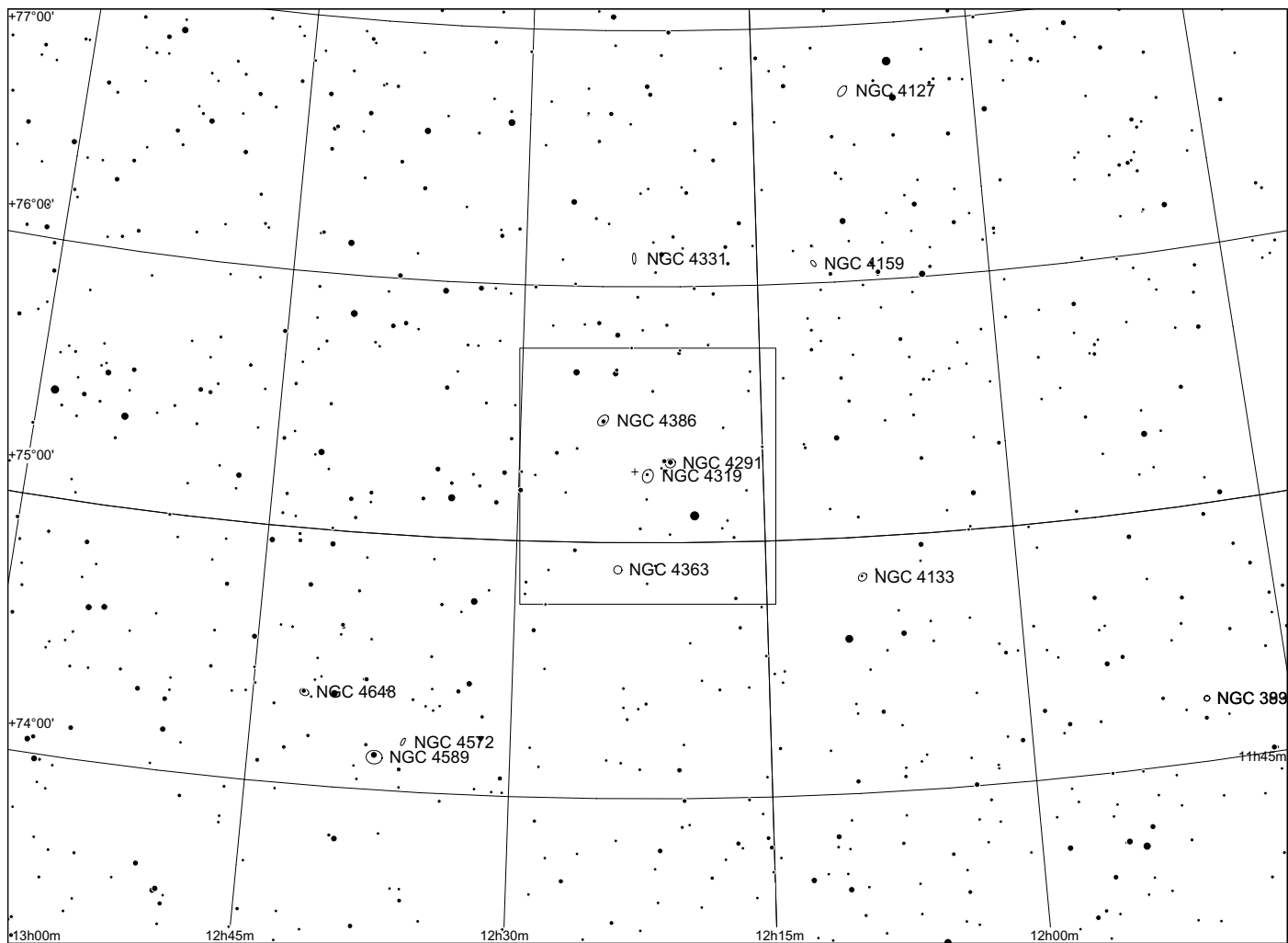
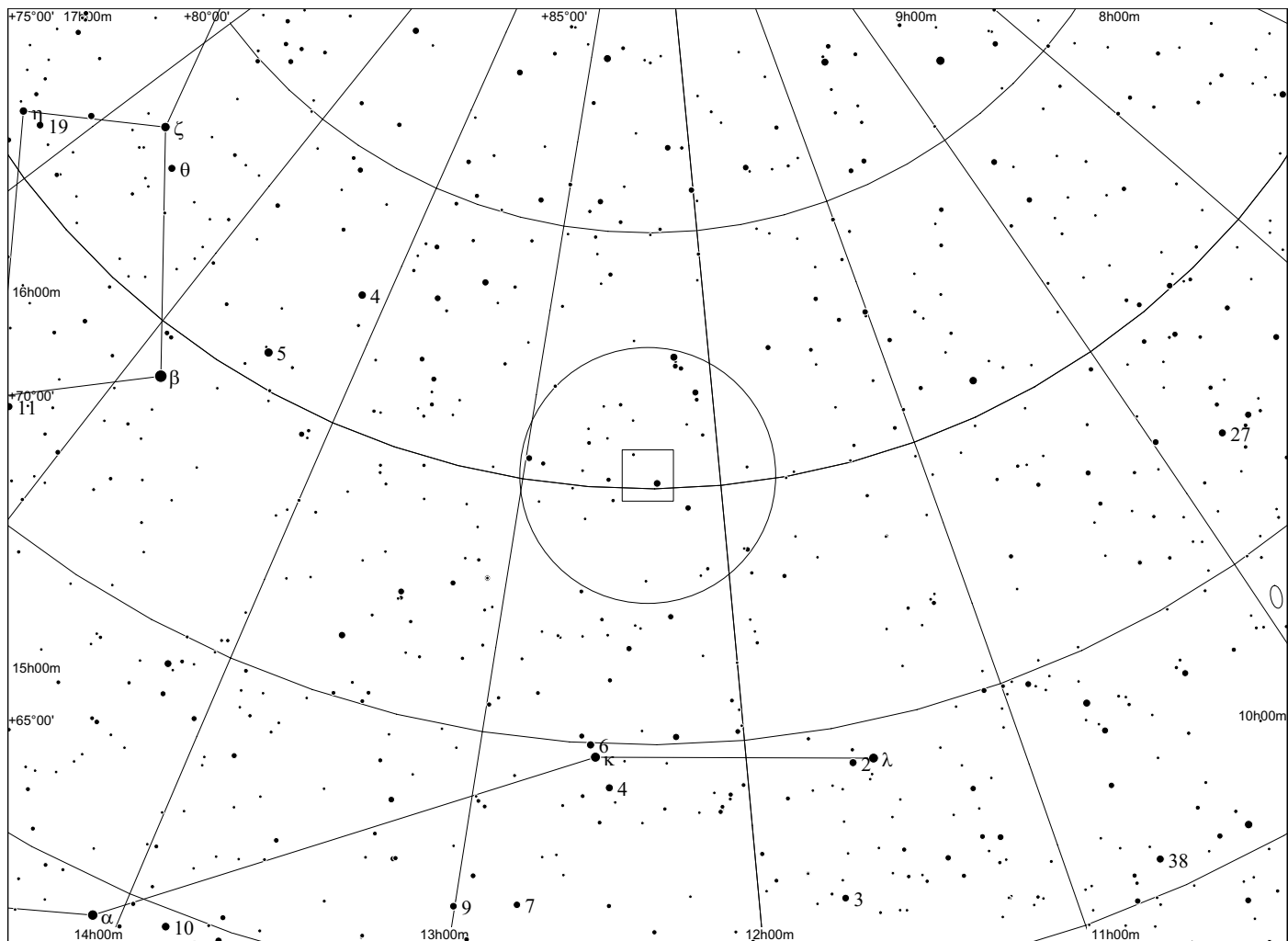
Markarian 205 in Draco



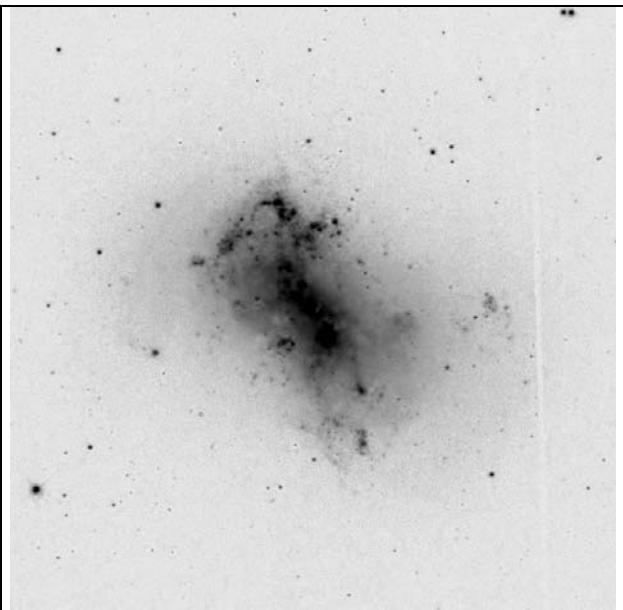
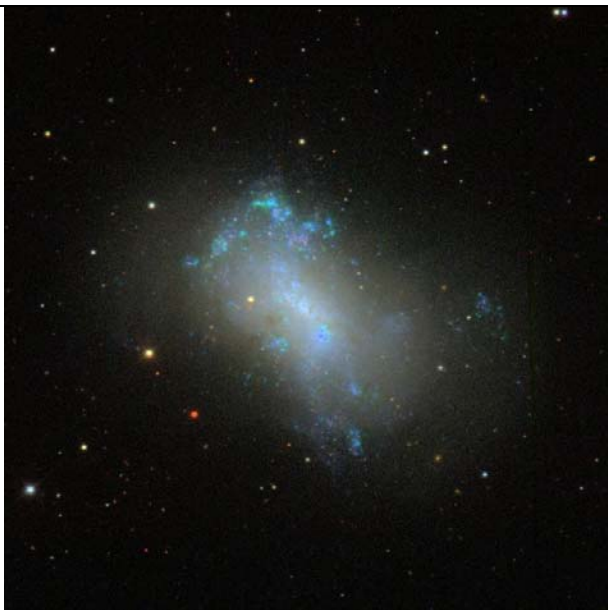
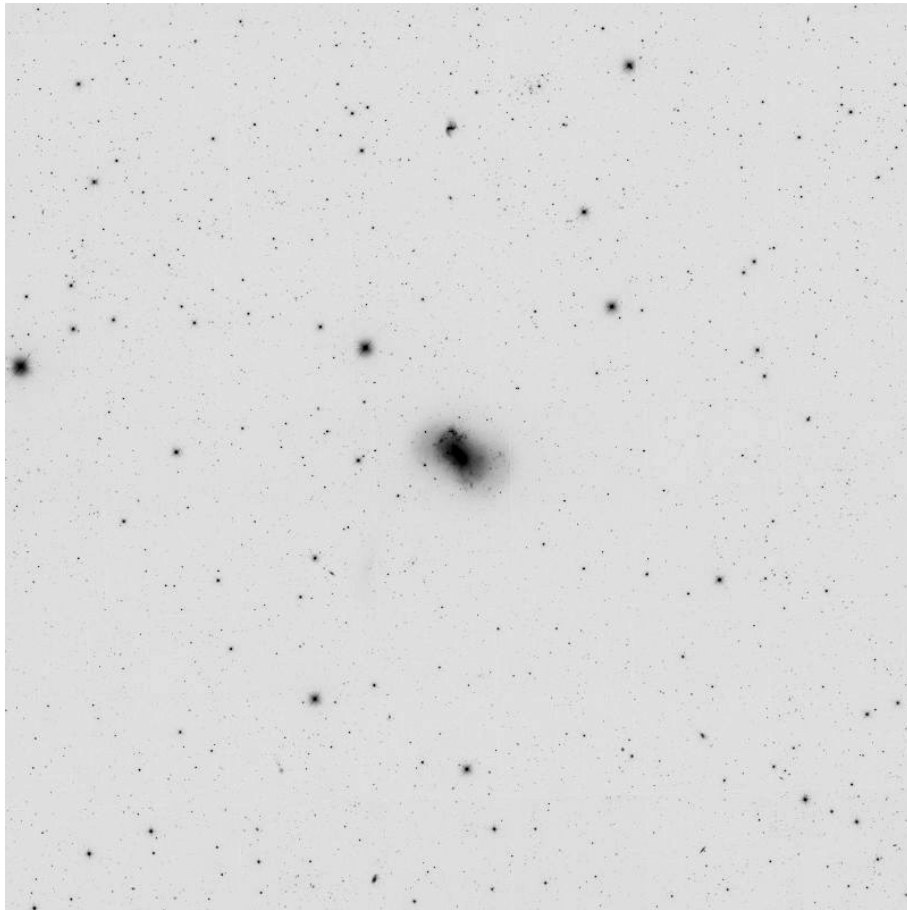
	other	RA	Dek	comments
Mrk 205	bei NGC4319	12 21 43	+75 19 21	Quasar

Mrk 205 ist ein geschichtsträchtiges Objekt. Trotz der extrem unterschiedlichen Rotverschiebungen von Mrk 205 und NGC4319 vermutete Halton Arp eine "Materiebrücke" zwischen beiden. Diese hätte bedeutet, dass beide Objekte nahe beieinander stehen müssen. Arp verfolgte die Hypothese, dass die hohe Rotverschiebung der Quasare nicht auf ihrer Entfernung beruhte, sondern dass sie aus Galaxien ausgestoßene Objekte sind und die Rotverschiebung ihrer Spektren andere Ursachen habe (z.B. Gravitation). Damit stand er im krassen Widerspruch zur damaligen wissenschaftlichen Community, allen voran Allan Sandage, was zu einigen Fehden zwischen den Anhängern der verschiedenen Sichtweisen führte.

Mrk 205 ist nicht einfach, ist aber schon mit 14" gut indirekt zu sehen.

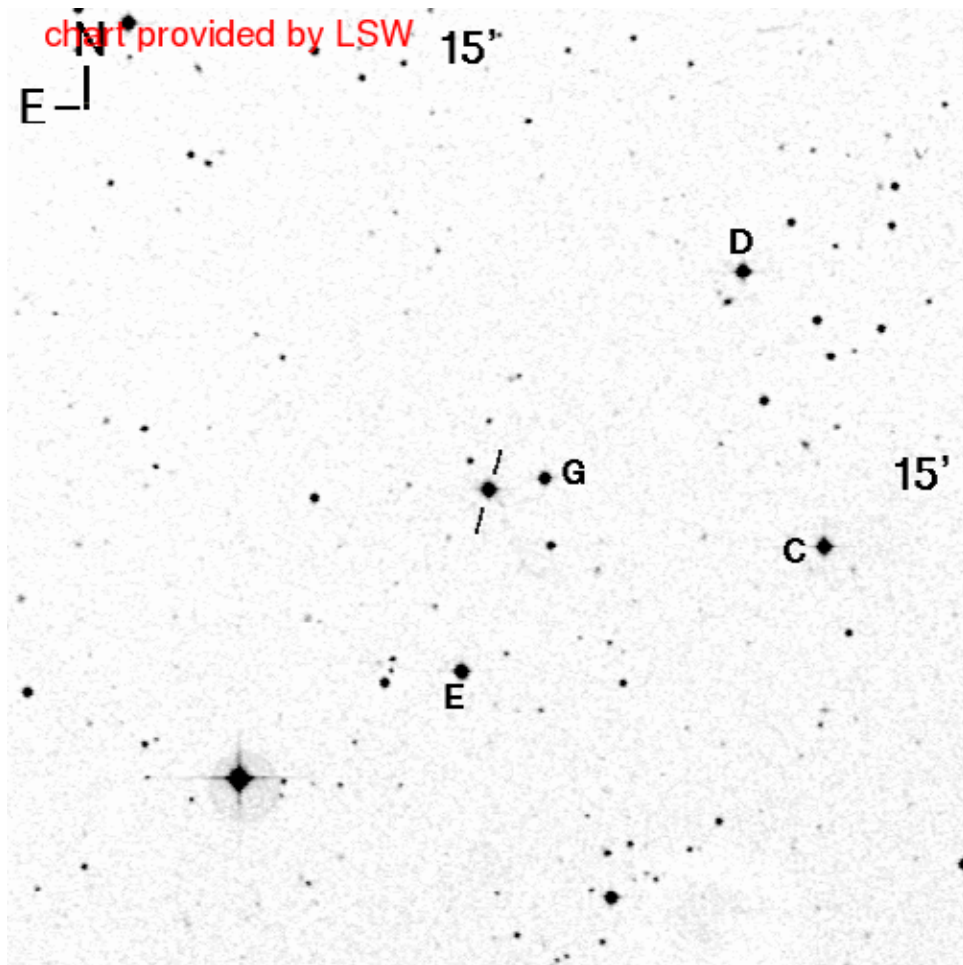


NGC 4449 in Canes Venatici

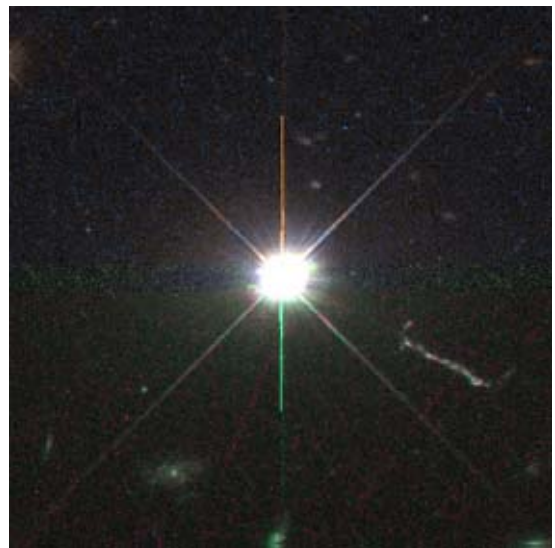


NGC 4449	12 28 11 +44 05 33	
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3C 273 in Virgo

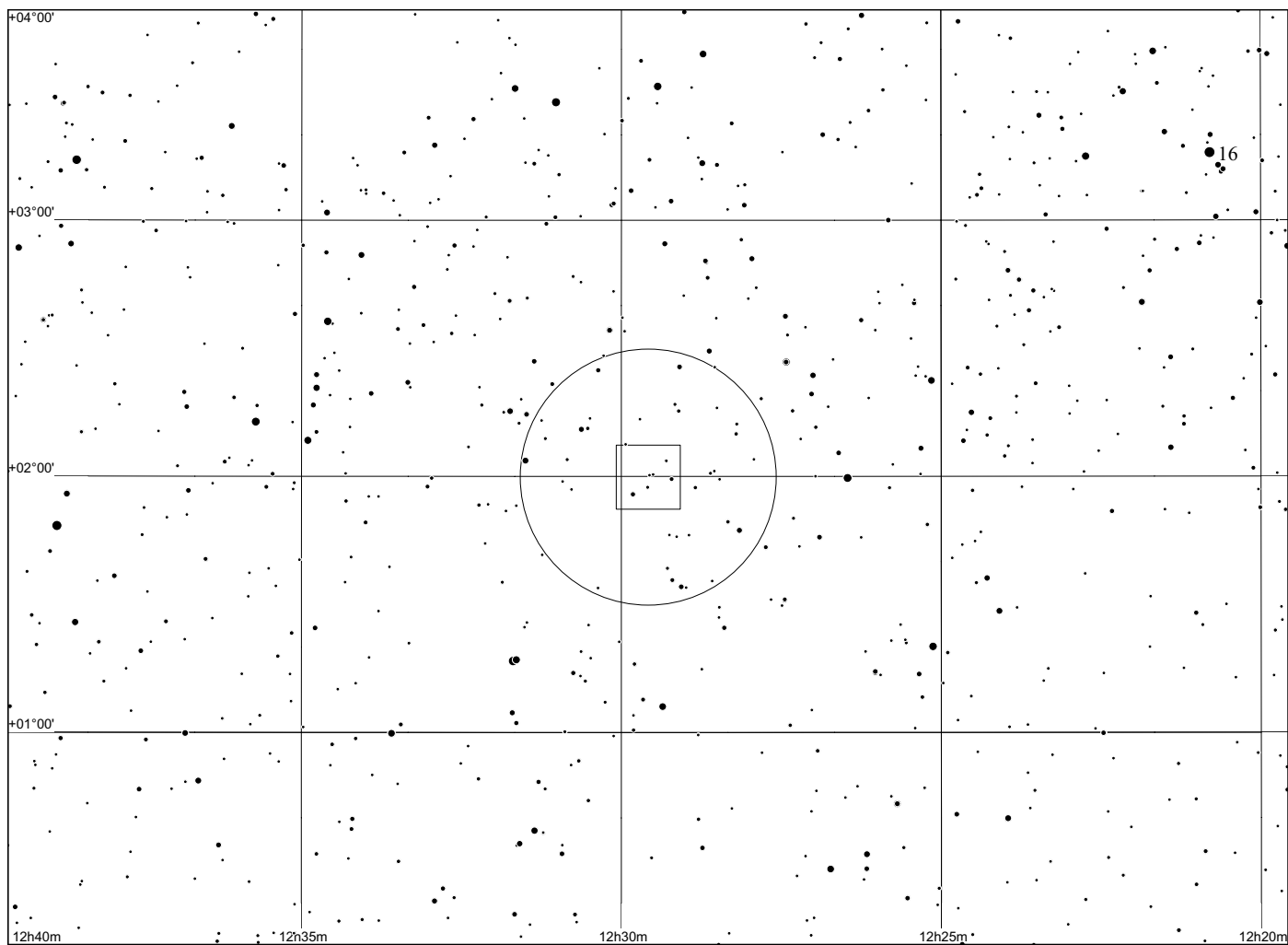
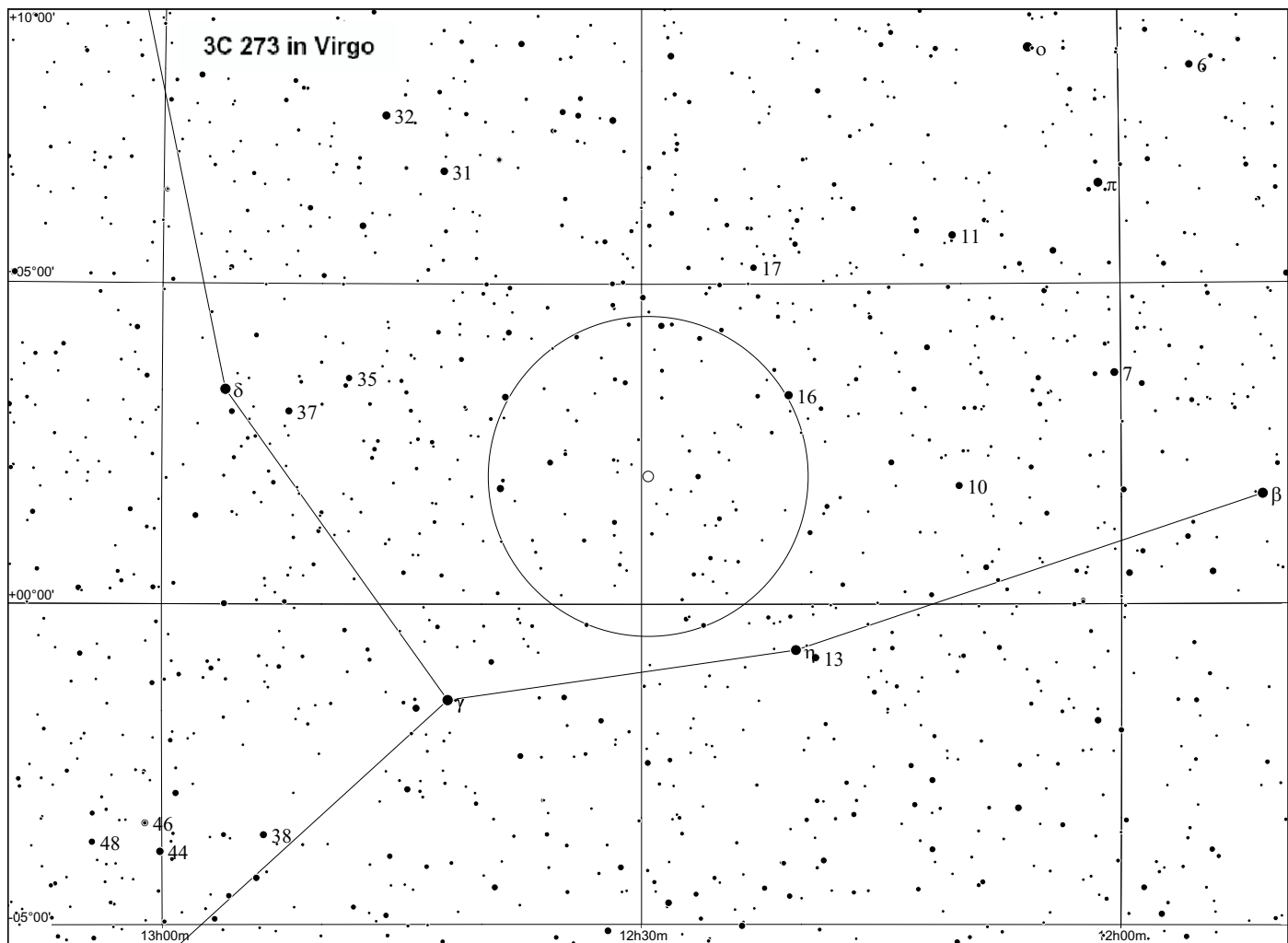


star	B	V	R
C	12.85(0.05)	11.87(0.04)	11.30(0.04)
D	13.17(0.05)	12.68(0.04)	12.31(0.04)
E	13.33(0.07)	12.69(0.04)	12.27(0.05)
G	14.12(0.05)	13.56(0.05)	13.16(0.05)

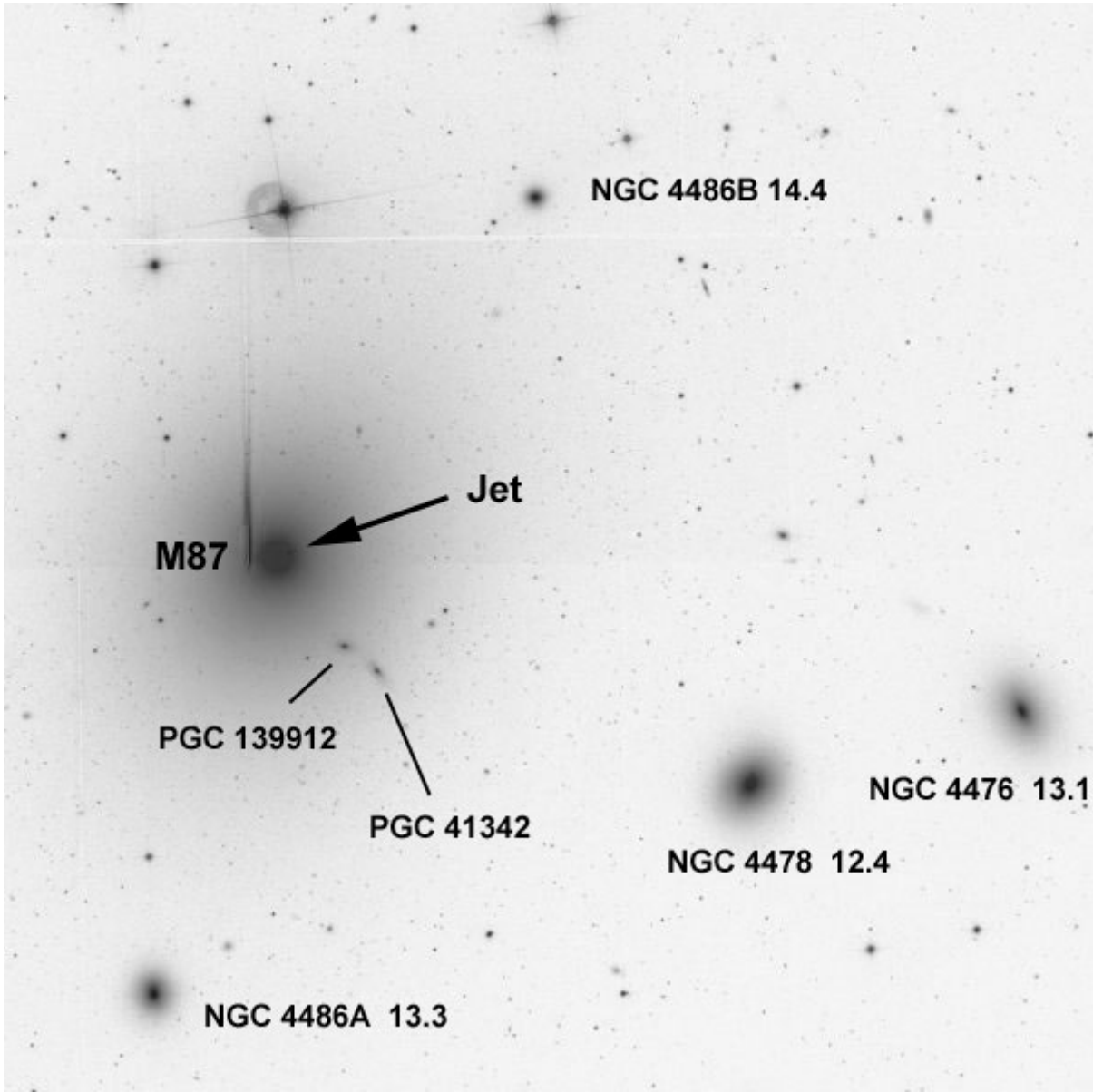


	other	RA	Dek	comments
3C 273		12 29 06	+02 03 08	first QSO

3C273 war der erste Quasar, bei dem das Spektrum von Maarten Schmidt als extrem rotverschoben erkannt wurde (was Allen Sandage, dem Nachfolger von Hubble an Palomar, ziemlich düpierte). Quasar sind Galaxien mit extrem hellen Kernen, in denen Materie auf eine Akkretionsscheibe um ein zentrales Schwarzes Loch stürzt. 3C273 hat eine Rotverschiebung von 0.158 und ist 2.4 Milliarden Lichtjahre entfernt.

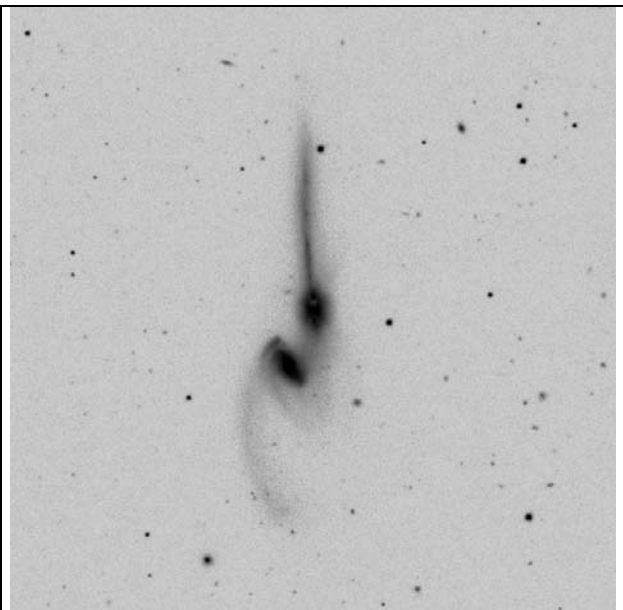
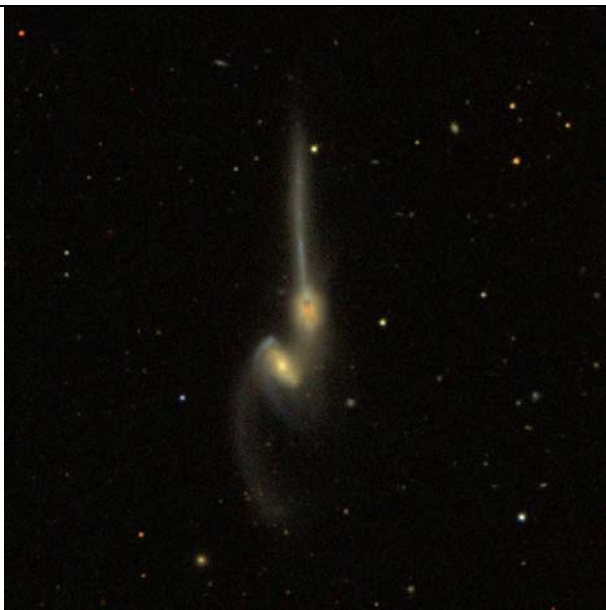
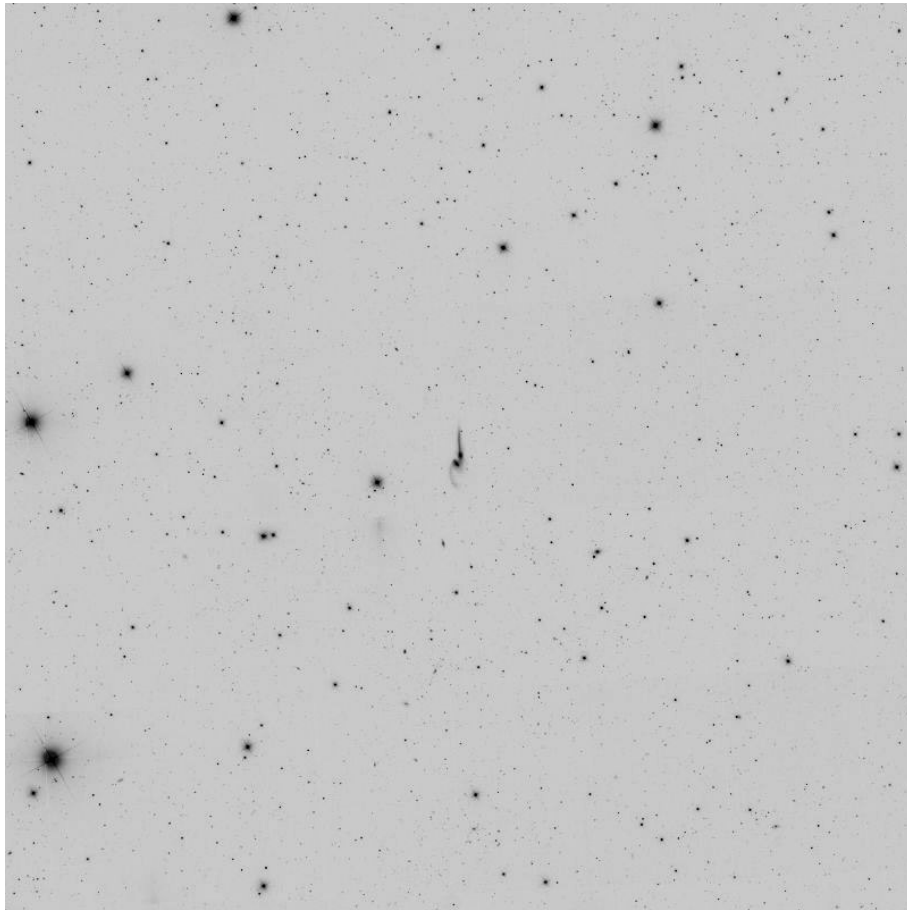


M87 in Virgo

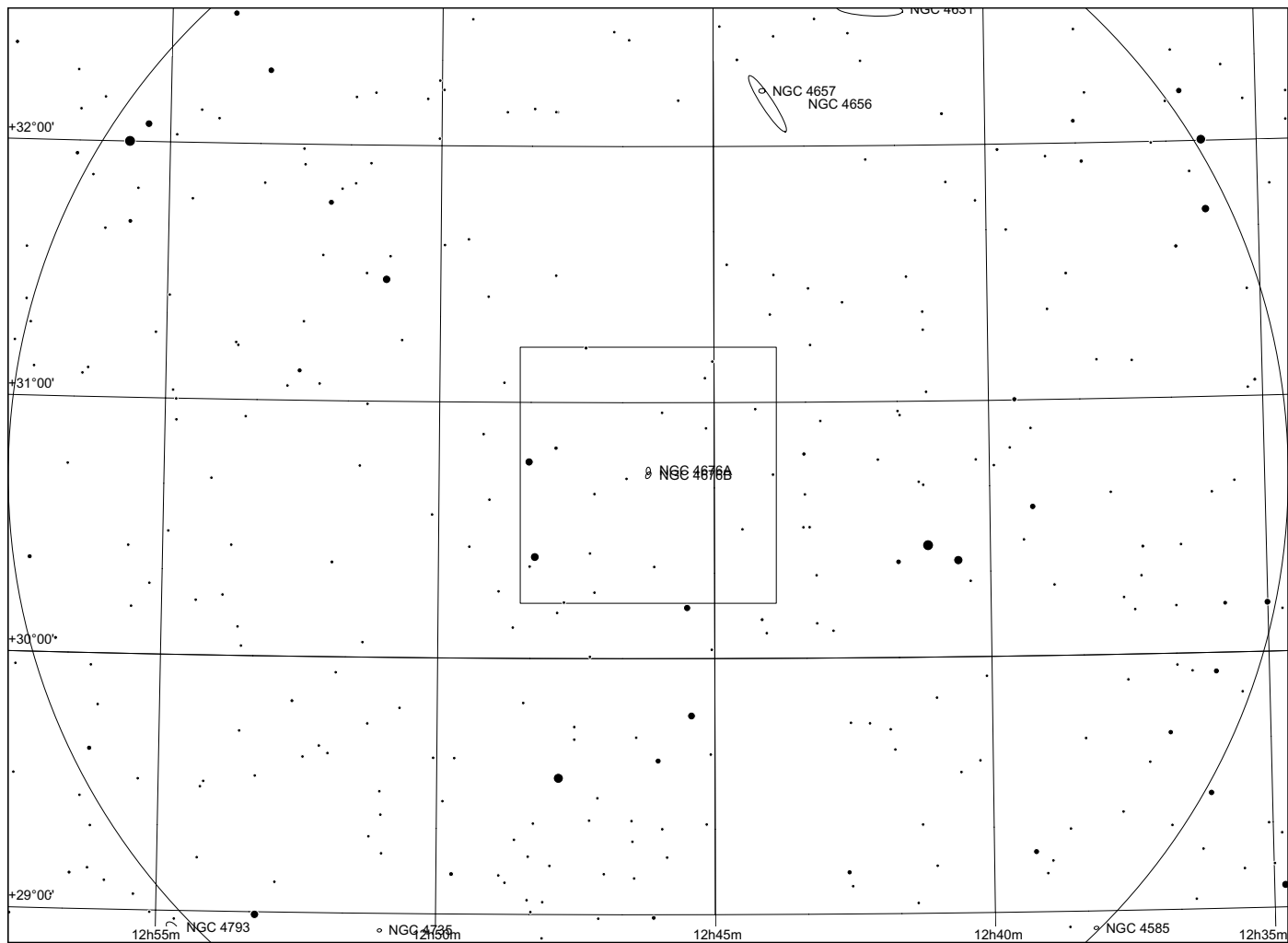
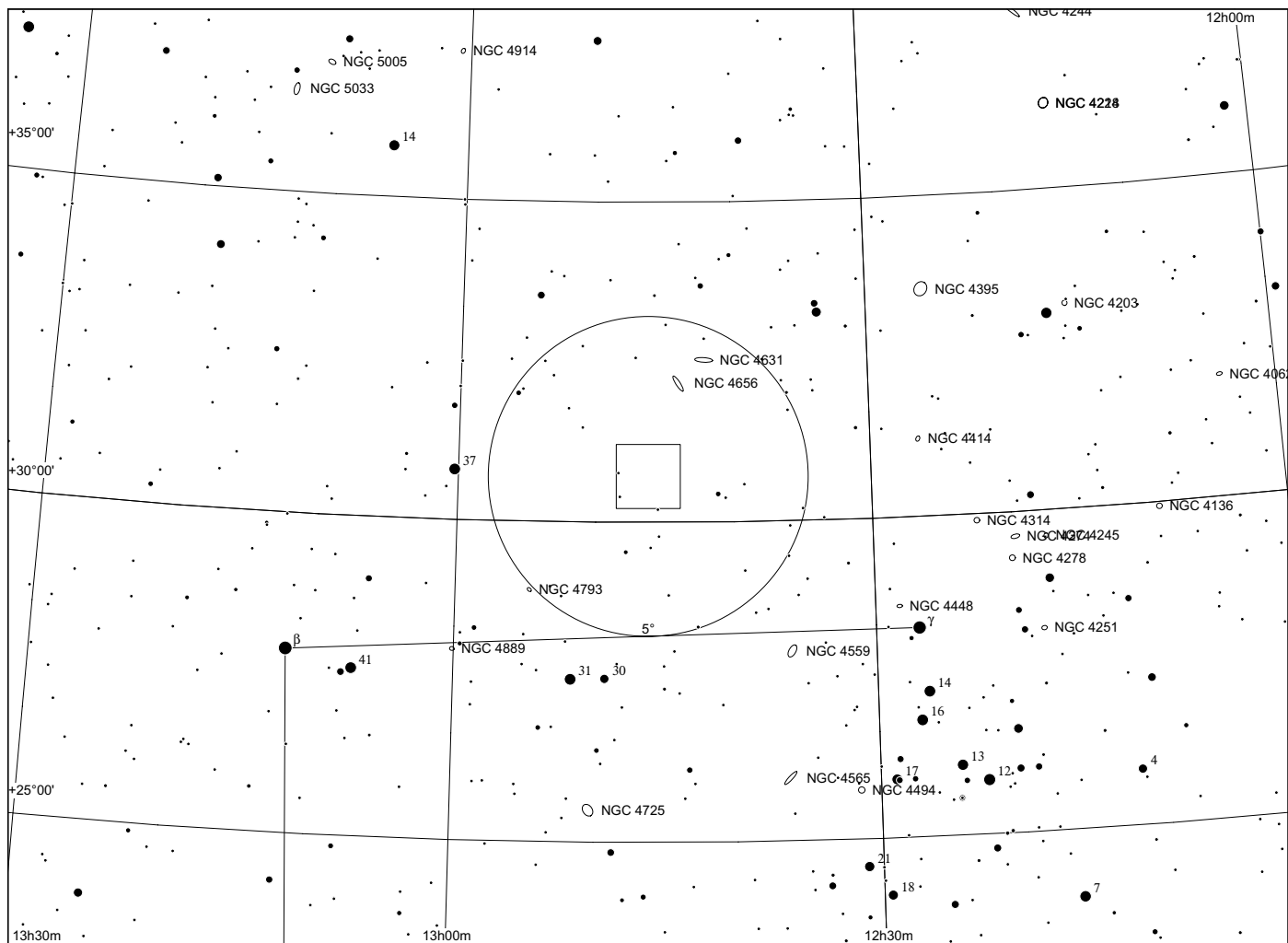


M87	12 30 49 +12 23 28	
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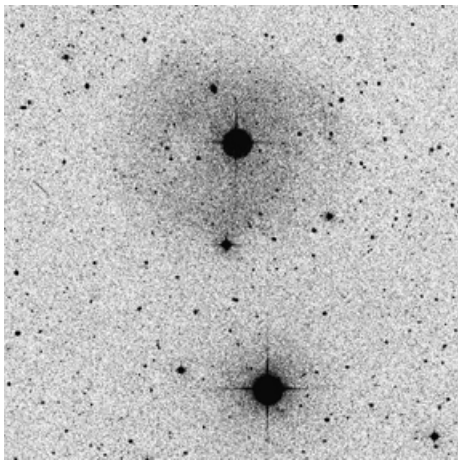
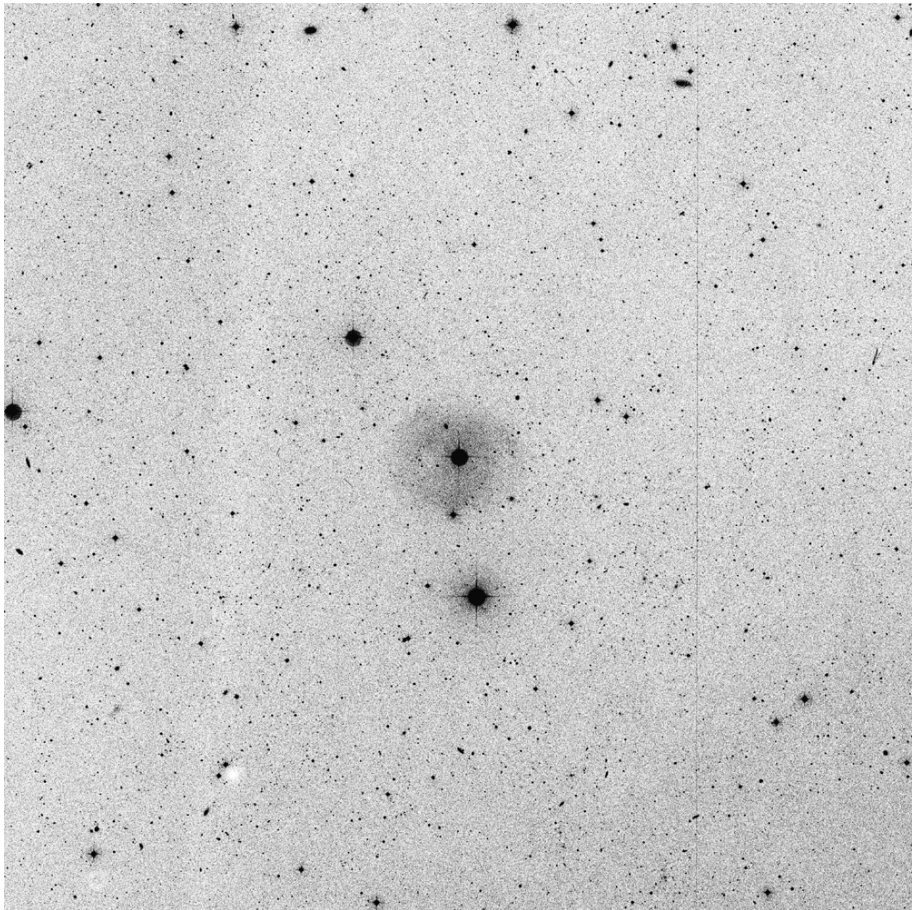
NGC 4676 in Coma Berenices



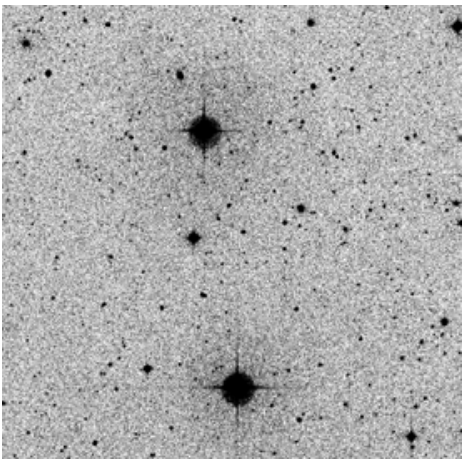
NGC 4676	12 46 00 +30 44 00	The Mice
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Longmore-Tritton (LoTr) 5 in Coma Berenices



POSS blue

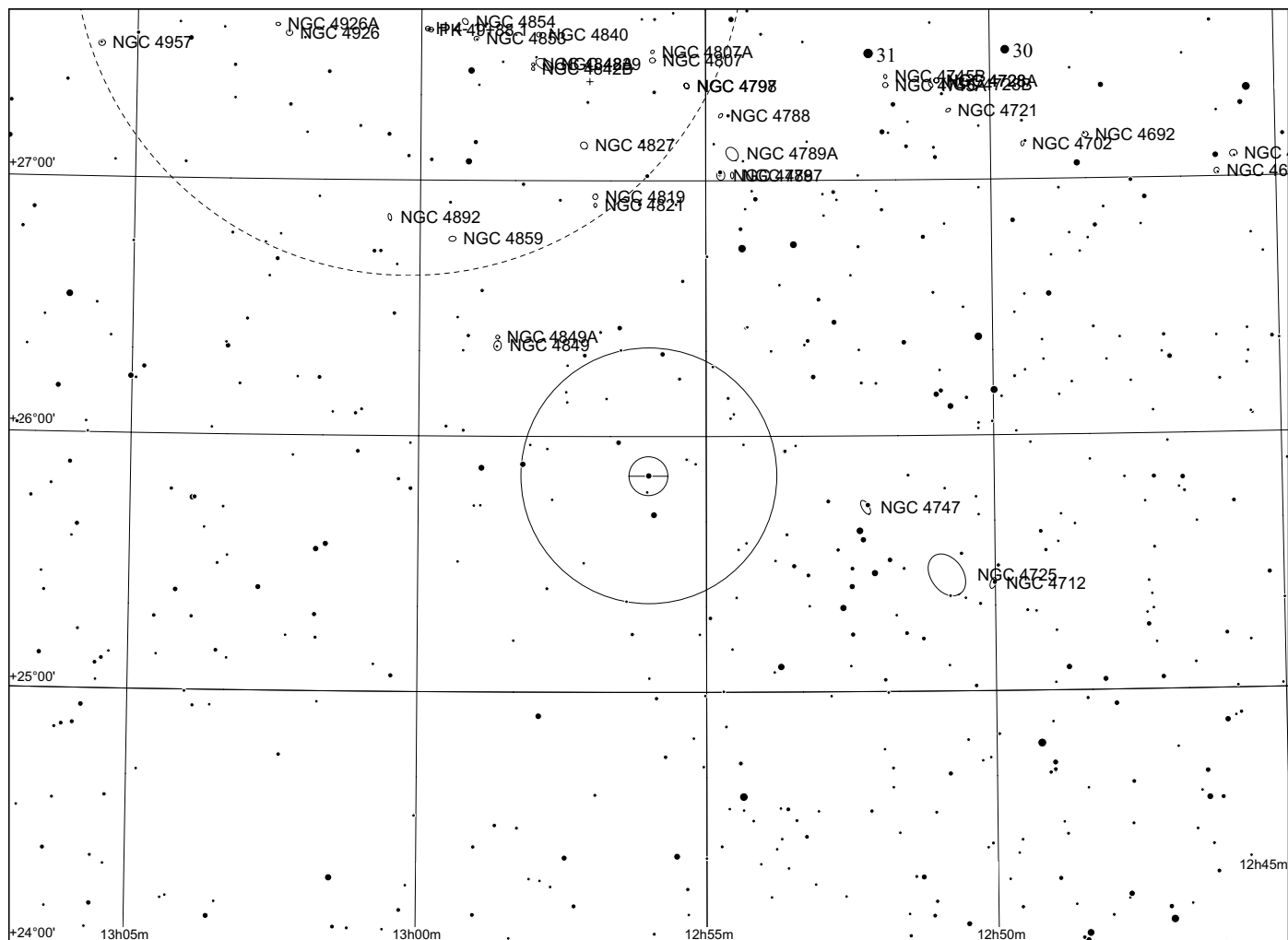
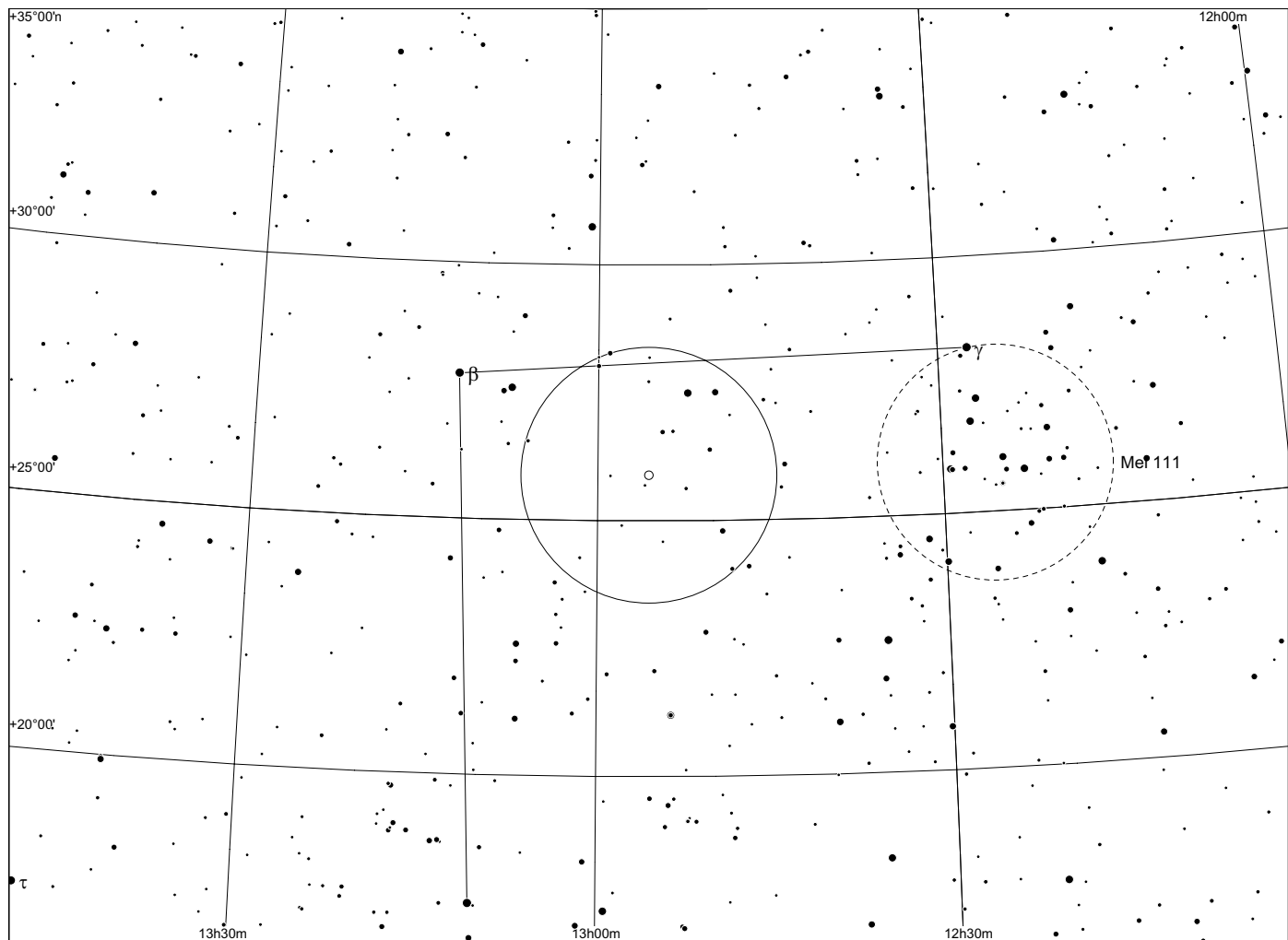


POSS red

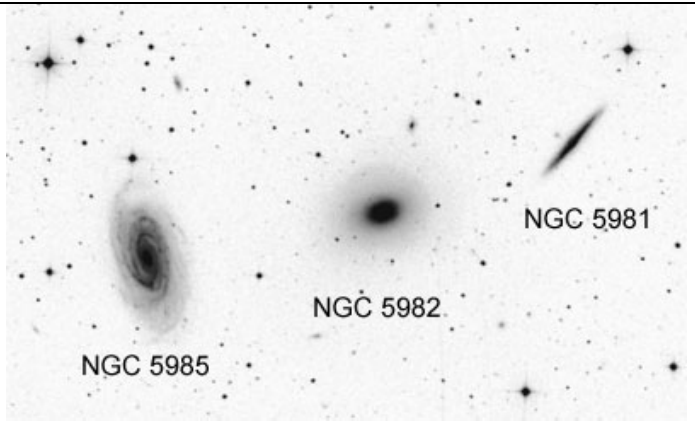
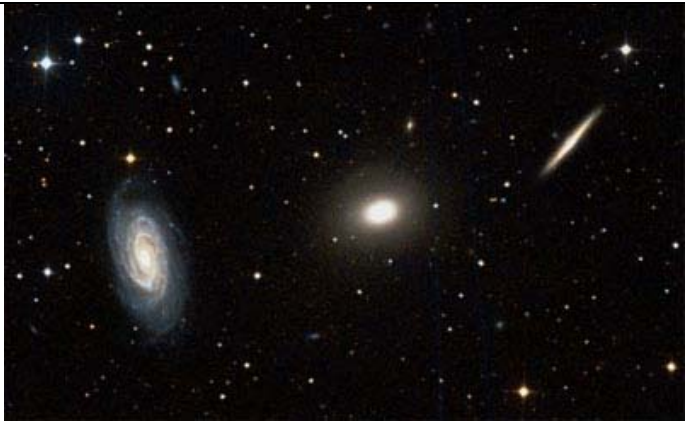
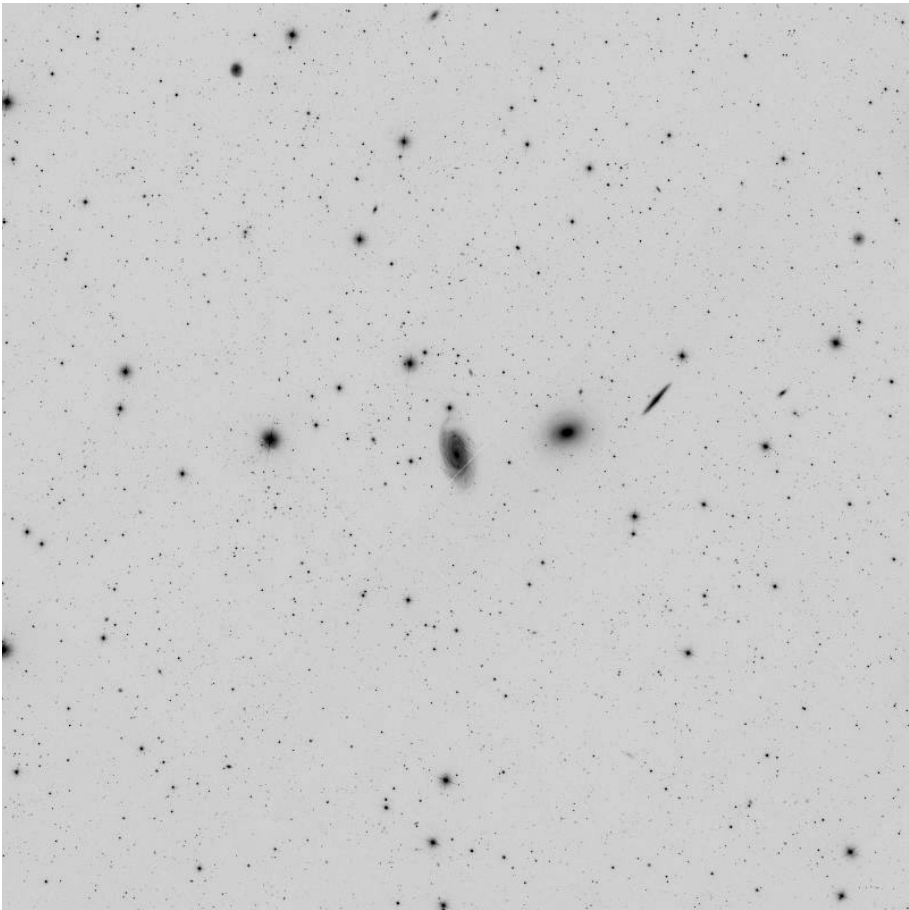
	RA	dec	diam [']	other	visibility on POSS II
LoTr 5	12h 55 34	+25 53 28	9* 8		good on blue, very faint on red

Observing notes:

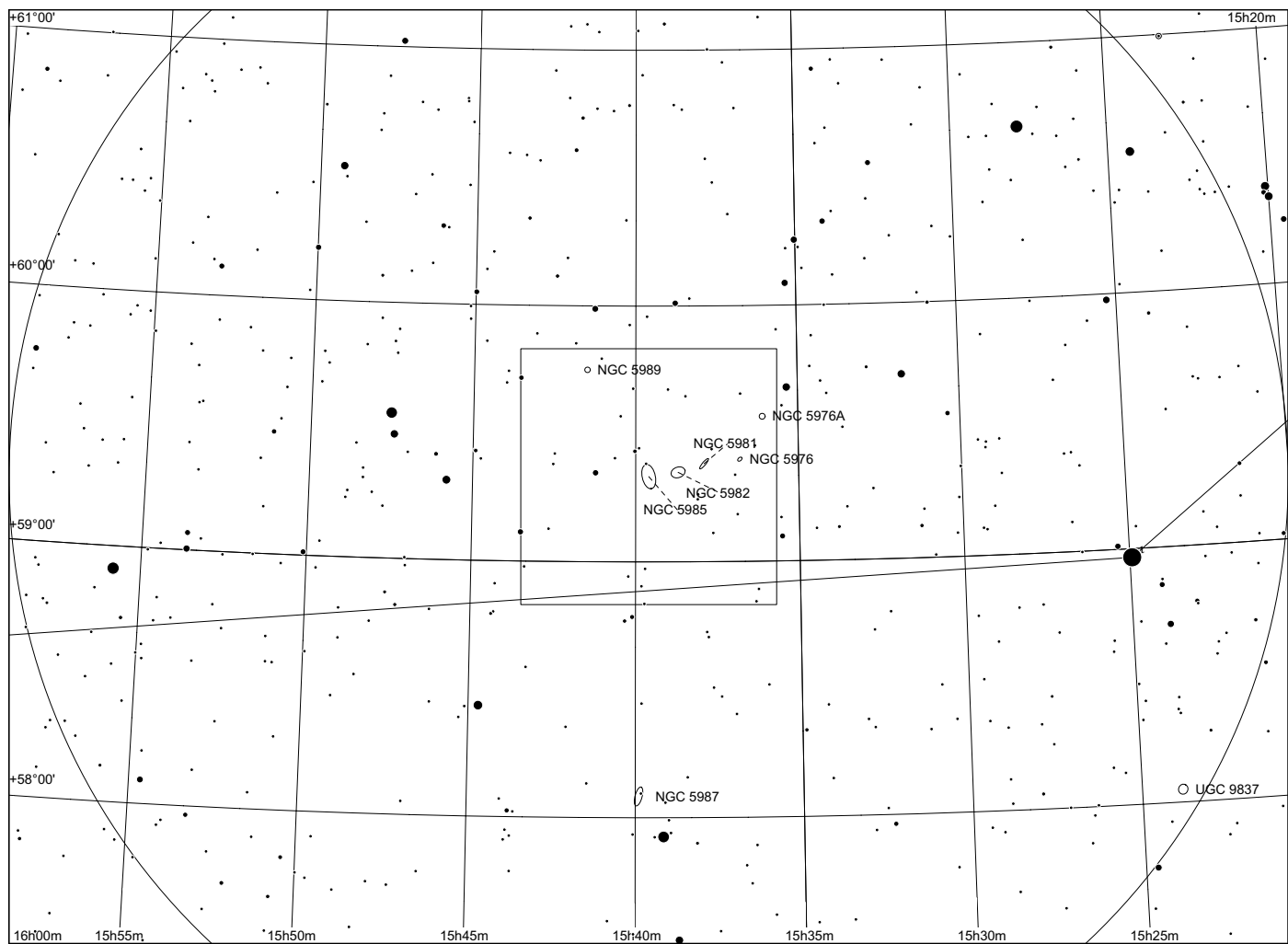
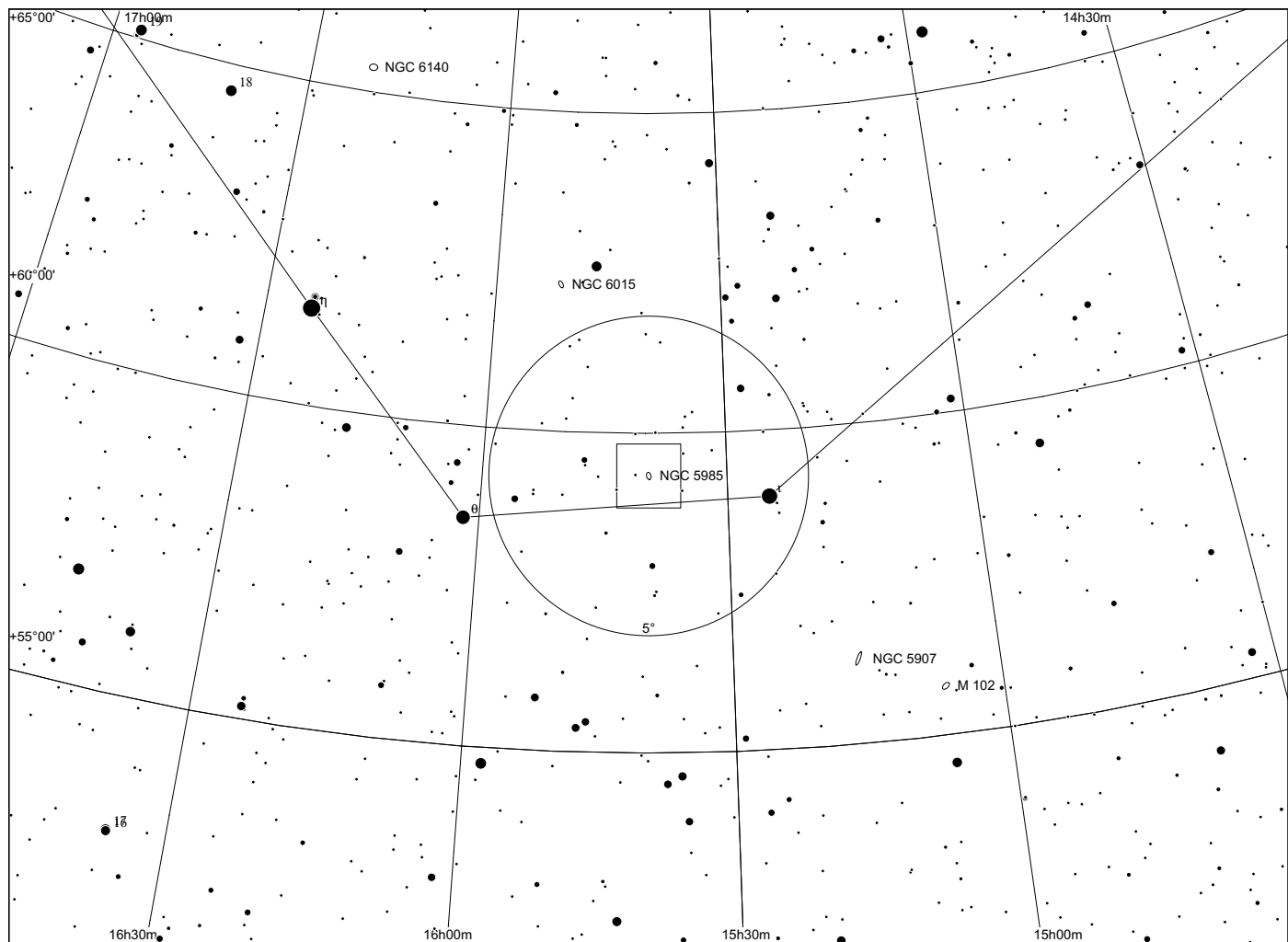
22" f/4.5
Successfully observed several times with OIII filtering. Not overly difficult. LoTr 5 shows a faint disk with relatively well defined boundary to the S, coinciding with the small star between the CS and the brighter star to the S. Also visible with UHC with somewhat lower contrast.



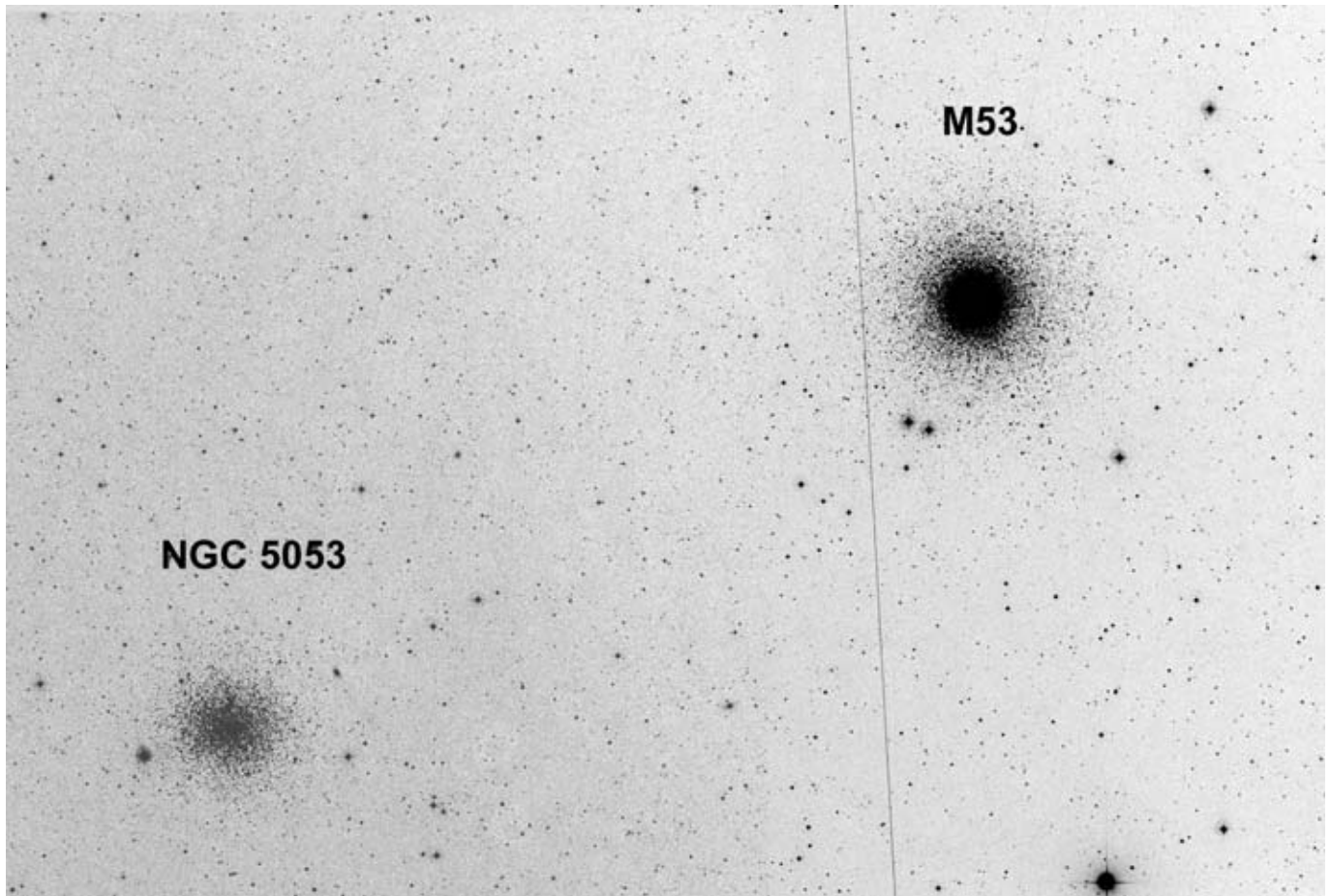
Draco Triplet in Draco



NGC 5985	13 08 12 +41 40 34	
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M53 / NGC5053 in Ophiuchus



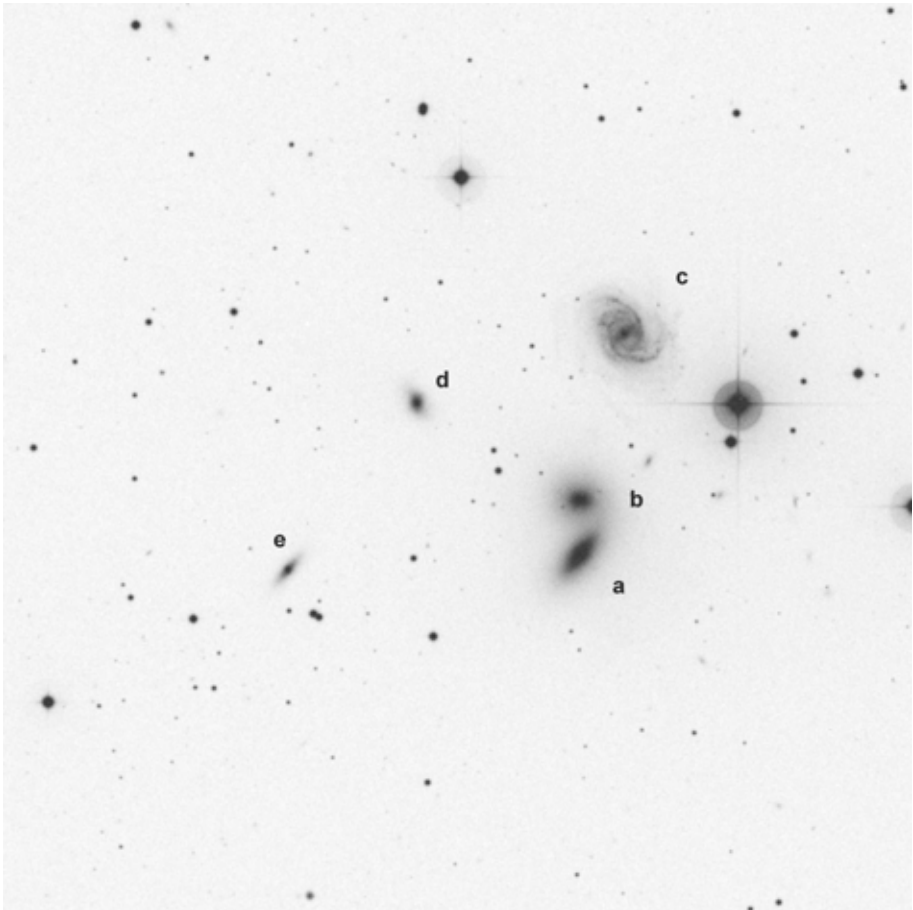
120x60 blue

	other	RA	Dek	comments
M53		13h 13	+18 10	class V
NGC5053				class XI

Zwei extrem unterschiedliche Kugelsternhaufen, die sehr eng beieinander stehen. Beide Kugelsternhaufen sind etwa 60 000 Lichtjahre entfernt, liegen also auch real nahe beieinander. Ihr unterschiedliches Erscheinungsbild beruht also auf ihrer unterschiedlichen Struktur und nicht auf unterschiedlicher Entfernung.

M53 ist ein schon recht stark konzentriert (Klasse V in der von I (stark konzentriert) bis XII (locker) reichenden Klassifizierung von Kugelsternhaufen), wahren 5053 nur Klasse XI erreicht. 5053 ist erheblich schwieriger als er auf der POSS Aufnahme erscheint! Er hat 9m8, besitzt etwa 3500 Sterne, von denen die hellsten 13m8 erreichen. "Normale" Kugelsternhaufen besitzen mehrere 100 000 Sterne.

Hickson 68 in Canes Venatici

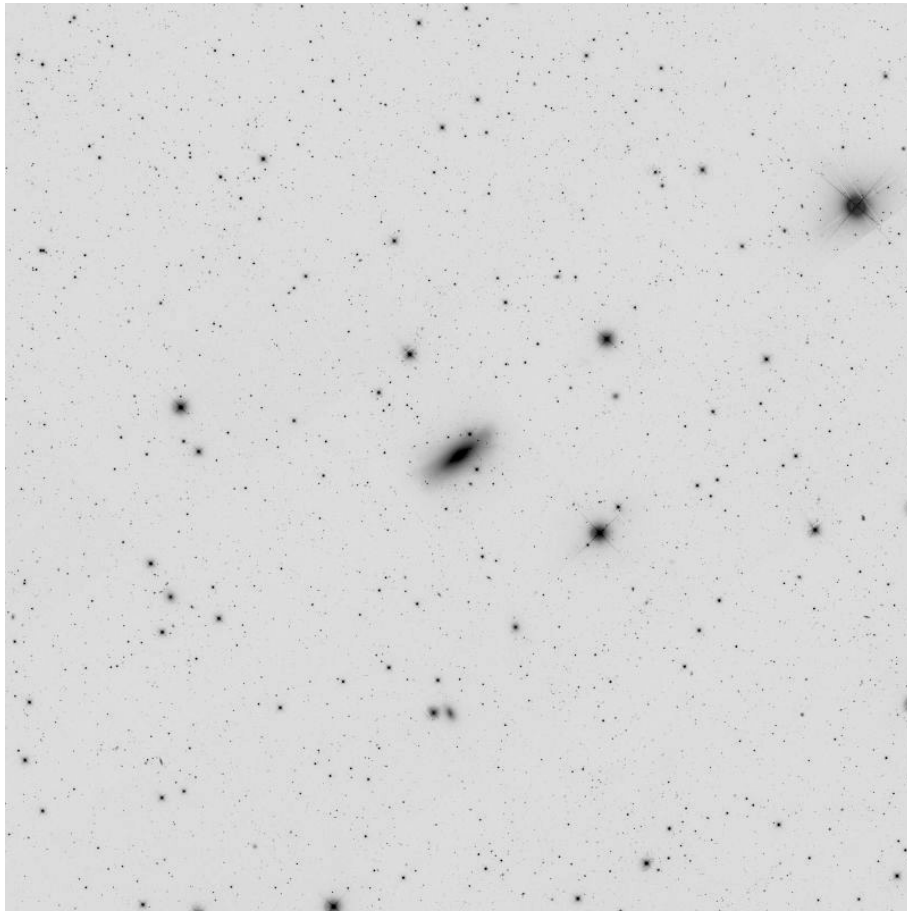


HCG Const. coordinates (2000) bright. memb. mag
68 CVn 13h 53m 41s +40° 20' NGC 5353 11.8

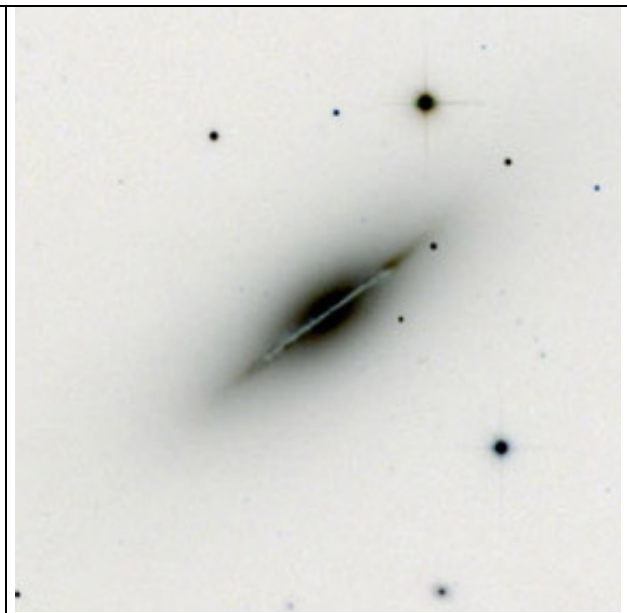
galaxy	ra (1950)	dec	a "	b "	type	T	B_I	C	D_B "	R_I	C	D_R "	B-R	B_T	B_TC	err	v_r km/s	err km/s	C	names
68a	13 51 19.59	+40 31 42.5	77.40	59.90	S0	1	12.04	1	112.30	10.19	1	197.30	1.63	12.08	11.84	0.08	2162	27	0	N5353,
U8813																				
68b	13 51 19.70	+40 32 52.8	86.40	71.10	E2	0	12.52	1	114.30	10.56	1	234.10	1.63	12.44	12.24	0.06	2635	23	0	N5354,
U8814																				
68c	13 51 14.81	+40 36 32.0	76.70	61.90	SBbc	6	12.11	2	134.30	10.90	2	166.40	1.10	12.20	11.93	0.20	2313	38	0	N5350,
Mk1485, U8810																				
68d	13 51 38.79	+40 35 00.5	31.10	21.80	E3	0	14.27	0	52.80	12.97	0	54.40	1.30	13.93	13.73	0.10	2408	29	0	N5355,
U8819																				
68e	13 51 52.87	+40 31 06.8	35.20	15.50	S0	1	14.74	0	43.70	13.27	0	56.40	1.40	14.56	14.22	0.10	2401	27	0	N5358,
U8826																				

Zusammen mit HCG 44 im Kopf des Löwen ist HCG 68 die hellste Hickson Gruppe mit e als schwierigster Einzelgalaxie. c besitzt ein Halo, das in großen Teleskopen strukturiert ist. NGC 5371 ist eine helle Spirale 30' ENE von HCG 68.

M102 in Draco

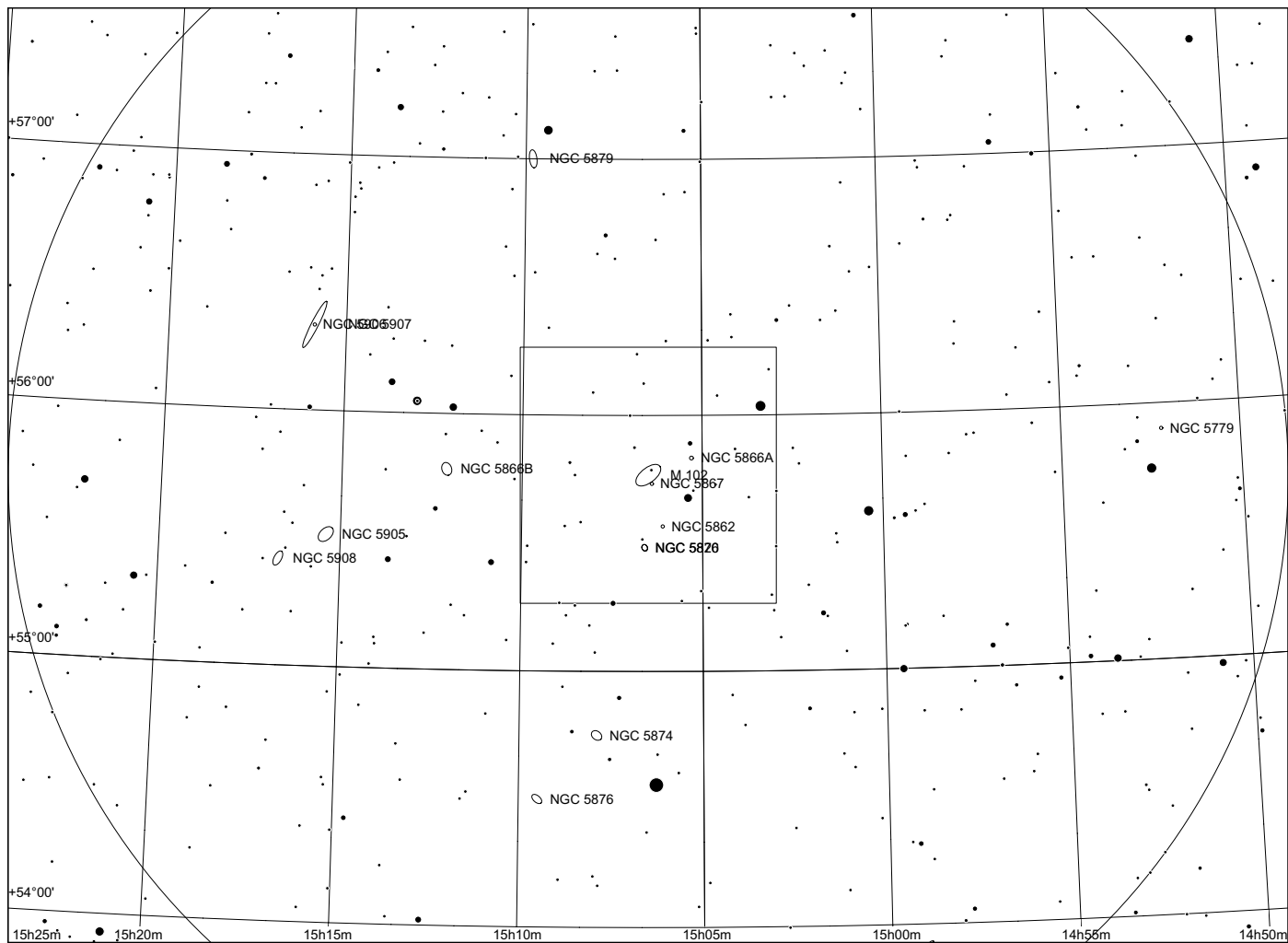
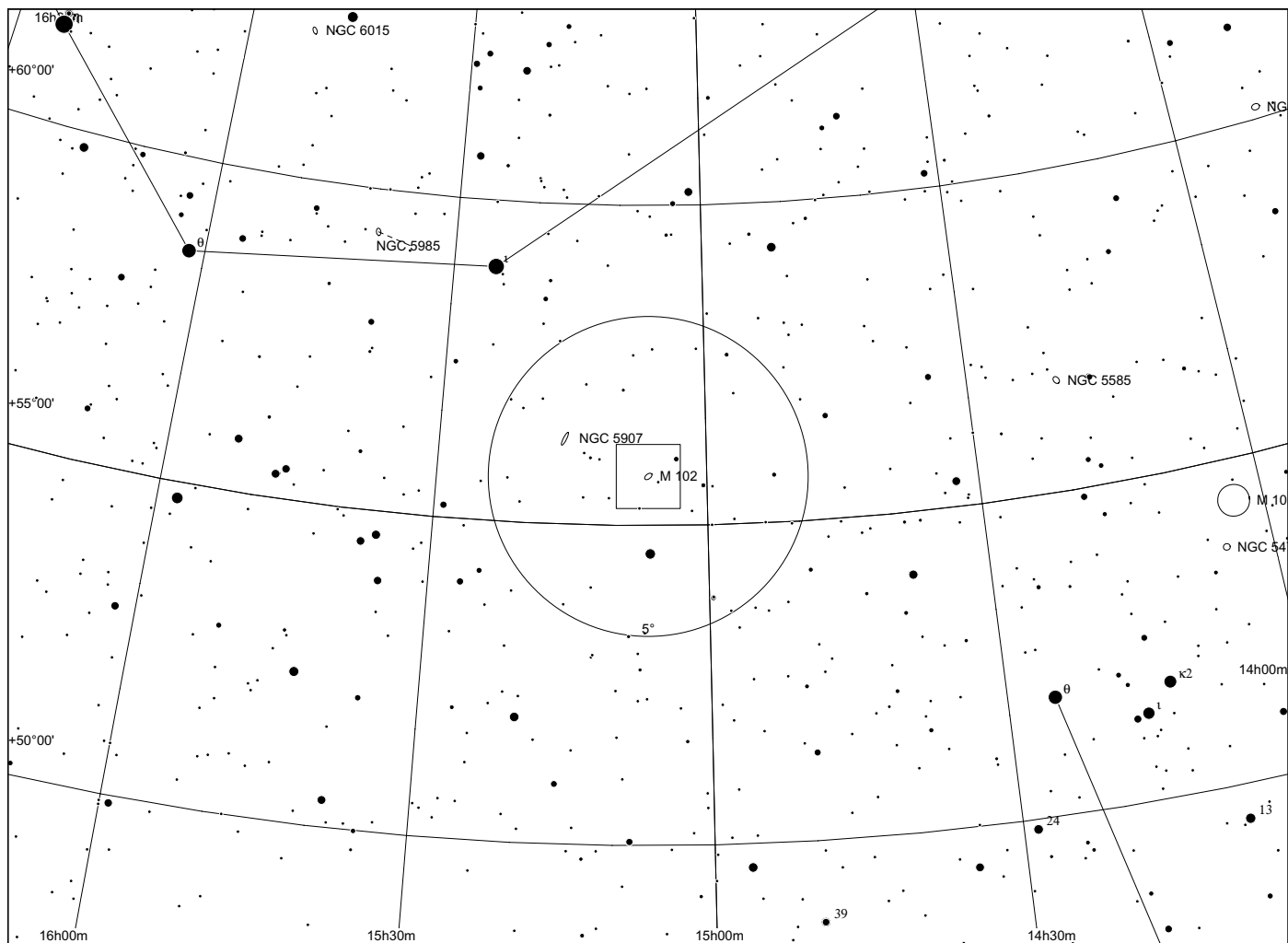


M102

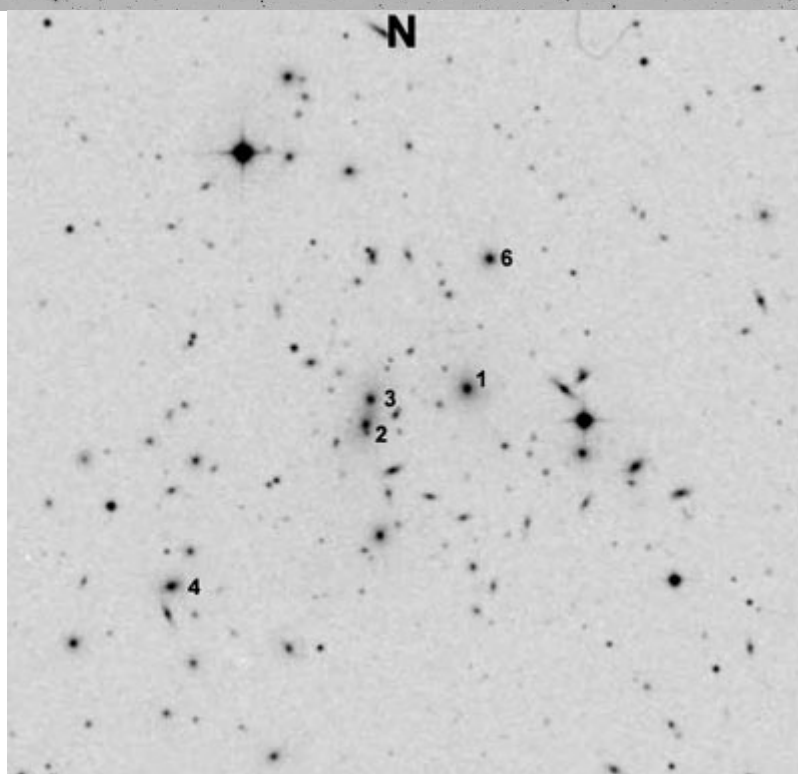
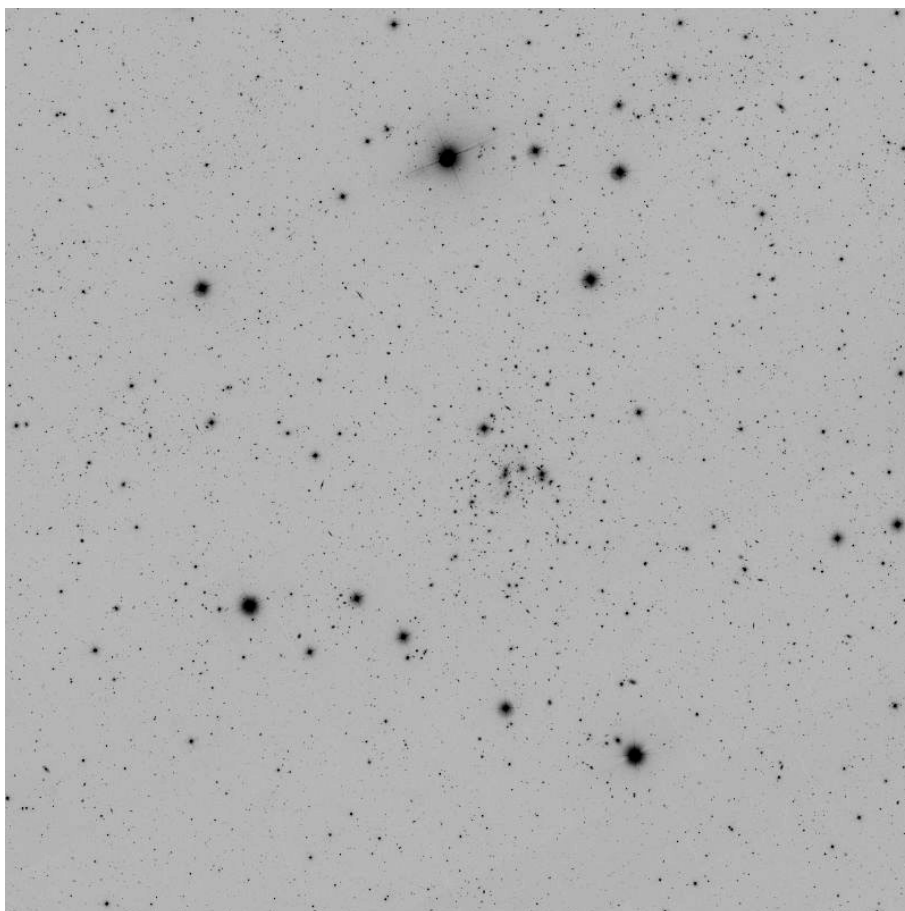


15 06 29 +55 45 47

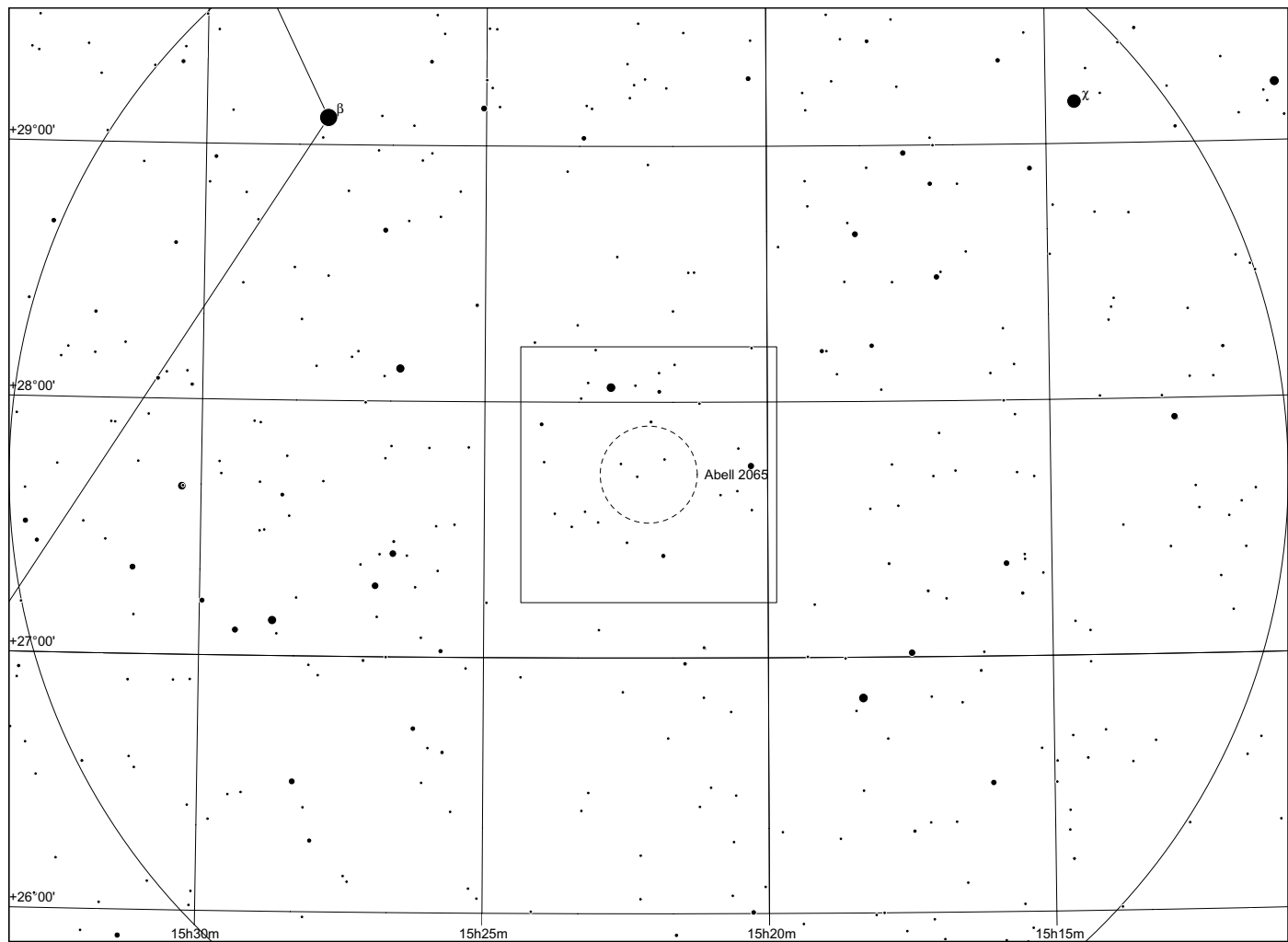
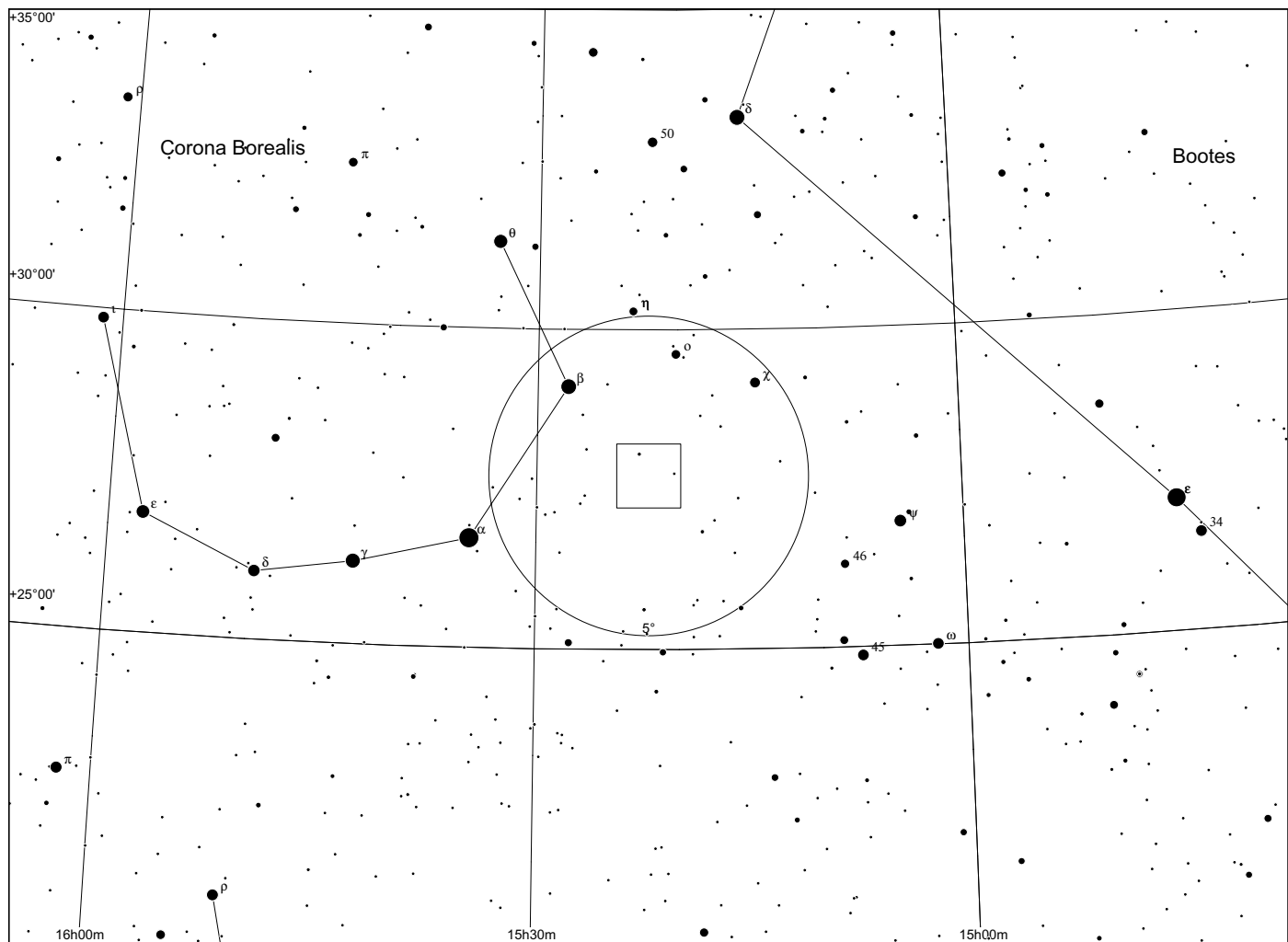
NGC 5866



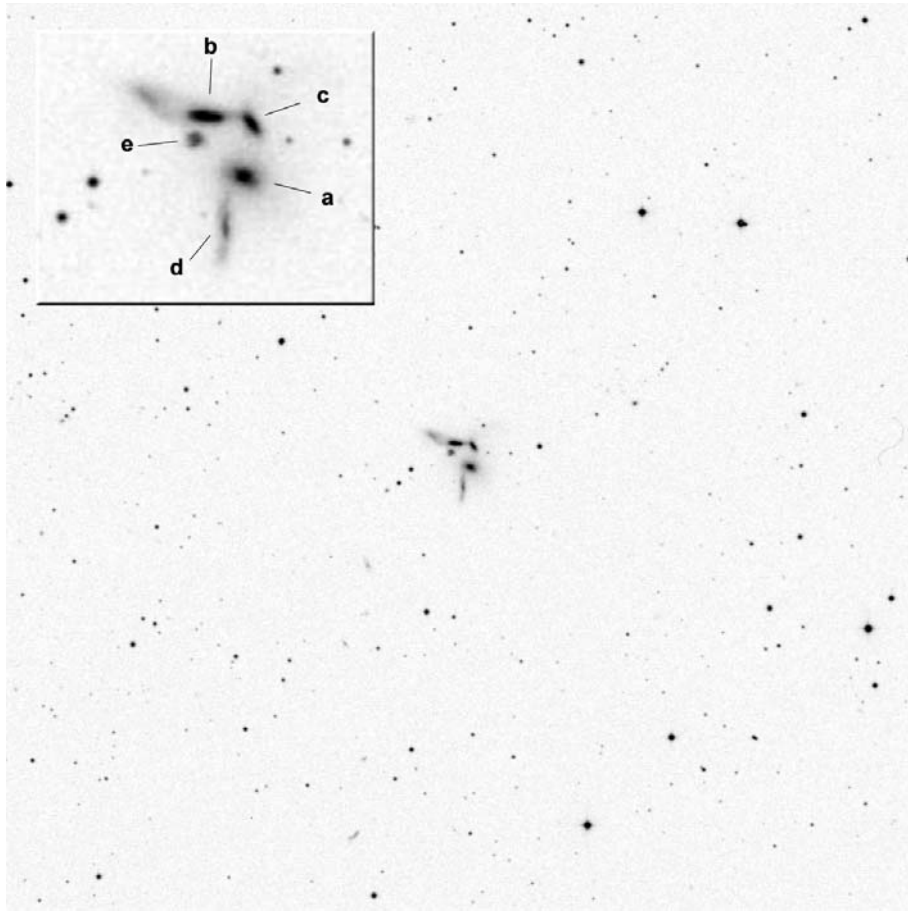
AGC 2065 in Corona Borealis



AGC 2065	15 22 42 +27 43 21	
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Hickson 79 in Serpens



HCG Const. coordinates (2000) bright. memb. mag
 79 Ser 15h 59m 13s +20° 45' NGC 6027 13.8 Seyfert's Sextet

galaxy	ra (1950)	dec	a "	b "	type	T	B_I	C	D_B "	R_I	C	D_R "	B-R	B_T	B_TC	err	v_r km/s	err km/s	C	names
79a U10116	15 56 59.59	+20 53 43.2	50.20	48.60	E0	0	15.12	1	44.70	13.46	1	50.50	1.60	14.70	14.35	0.20	4292	35	0	N6027,
79b U10116	15 57 00.80	+20 54 15.4	60.50	44.10	S0	1	14.55	1	43.00	13.07	1	46.90	1.44	14.19	13.78	0.20	4446	25	0	N6027,
79c U10116	15 56 59.14	+20 54 09.8	40.30	26.10	S0	1	15.48	1	33.80	14.21	1	35.70	1.27	15.15	14.72	0.20	4146	50	0	N6027,
79d U10116	15 57 00.18	+20 53 15.5	28.10	12.30	Sdm	10	16.89	1	23.50	16.74	1	16.10	0.85	16.51	15.87	0.20	4503	43	0	N6027,
79e U10116	15 57 01.31	+20 54 01.5	11.10	8.60	Scd	8	16.61	1	25.10	15.31	1	26.50	1.32	16.31	15.87	0.20	19809	50	0	N6027,

Notes:

22" f/4: Famous Seyfert's sextet, being actually a quintet with the sixth member being probably a tidal plume of b. This is an extremely challenging group requiring aperture and high magnification. It is the prototype of the extremely tight members of Hickson's catalog.

HCG 79 appears as a horseshoe-shaped glow of about 2' diameter. The group can be separated intermittently into three individual knots (a, c, and b) using 360x, still it is difficult to pin down their precise location permanently.

